

# Solar Impact on Earth Climate

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- Source of energy for the Earth System /planets

What are the sources of variation in solar radiation?

- Solar activity
- Orbital variations

Do these changes affect the climate?

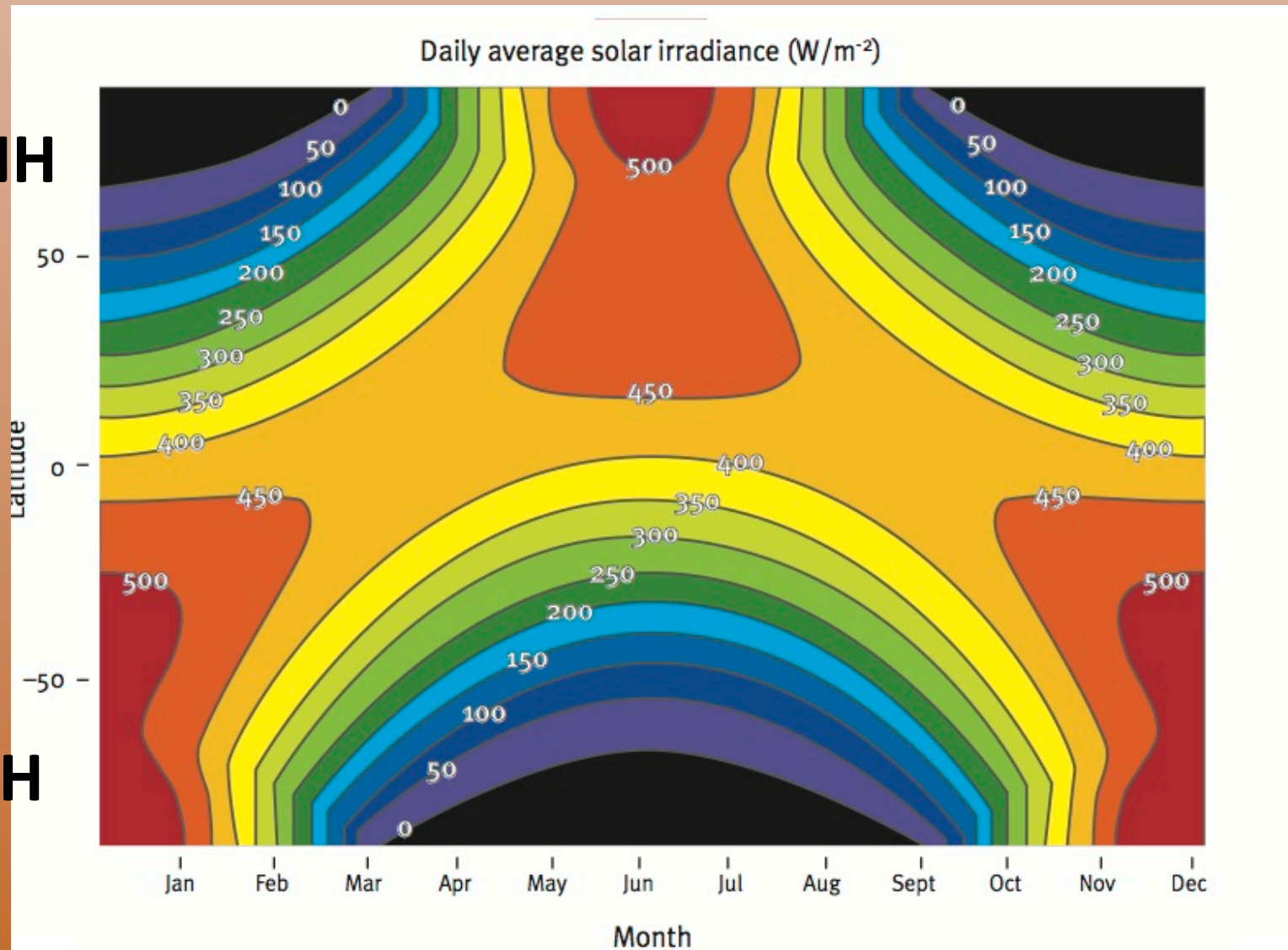
Do we understand the processes involved?

Gray et al, Rev Geophys, 2010 and Haigh et al.

# Distribution of Solar radiation at TOA

NH

SH



# Distribution of Solar radiation at TOA

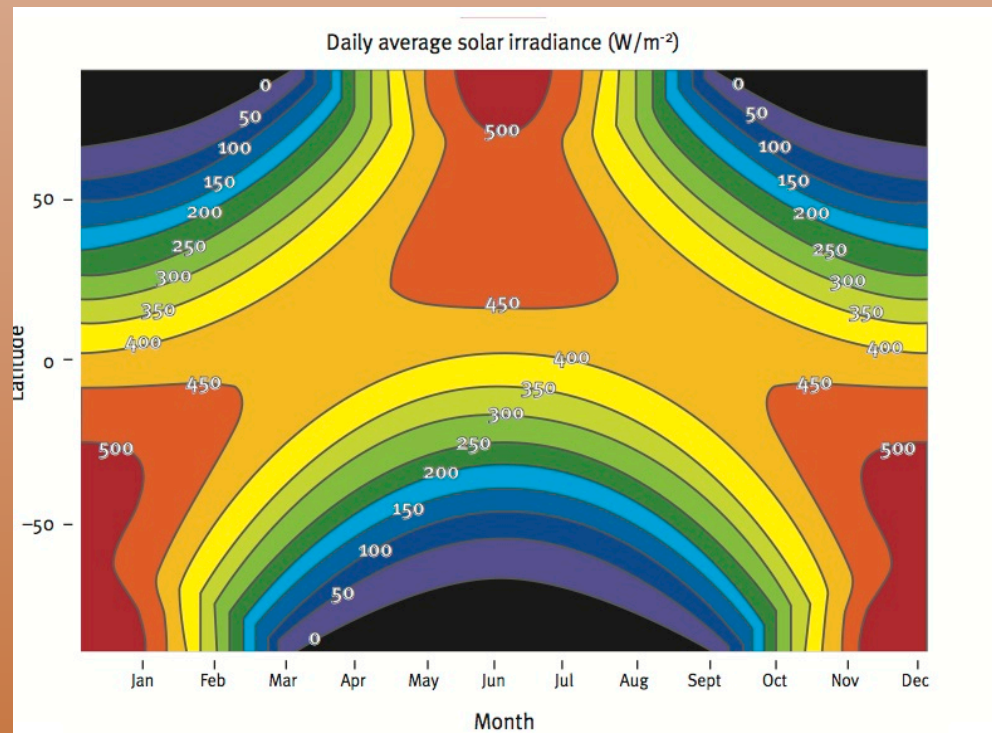
Mean E-S distance:  
 $1.496 \times 10^{11} \text{ m (= AU)}$

Max E-S distance:  
 $1.521 \times 10^{11} \text{ m (= 1.017 AU)}$

Min E-S distance:  
 $1.471 \times 10^{11} \text{ m (= 0.983 AU)}$

Eccentricity = 0.0167

Tilt =  $23^{\circ}27'$



# Geometry

$$F_{SW} = F_{SW}^0 \cos(Z)$$

$$F_{SW}^0 = S(d_m / d)^2$$

$$\cos Z = \sin \phi \sin \delta + \cos \phi \cos \delta \cosh$$

F = irradiance  
Z= zenith angle  
d=distance E-S  
h=hour angle

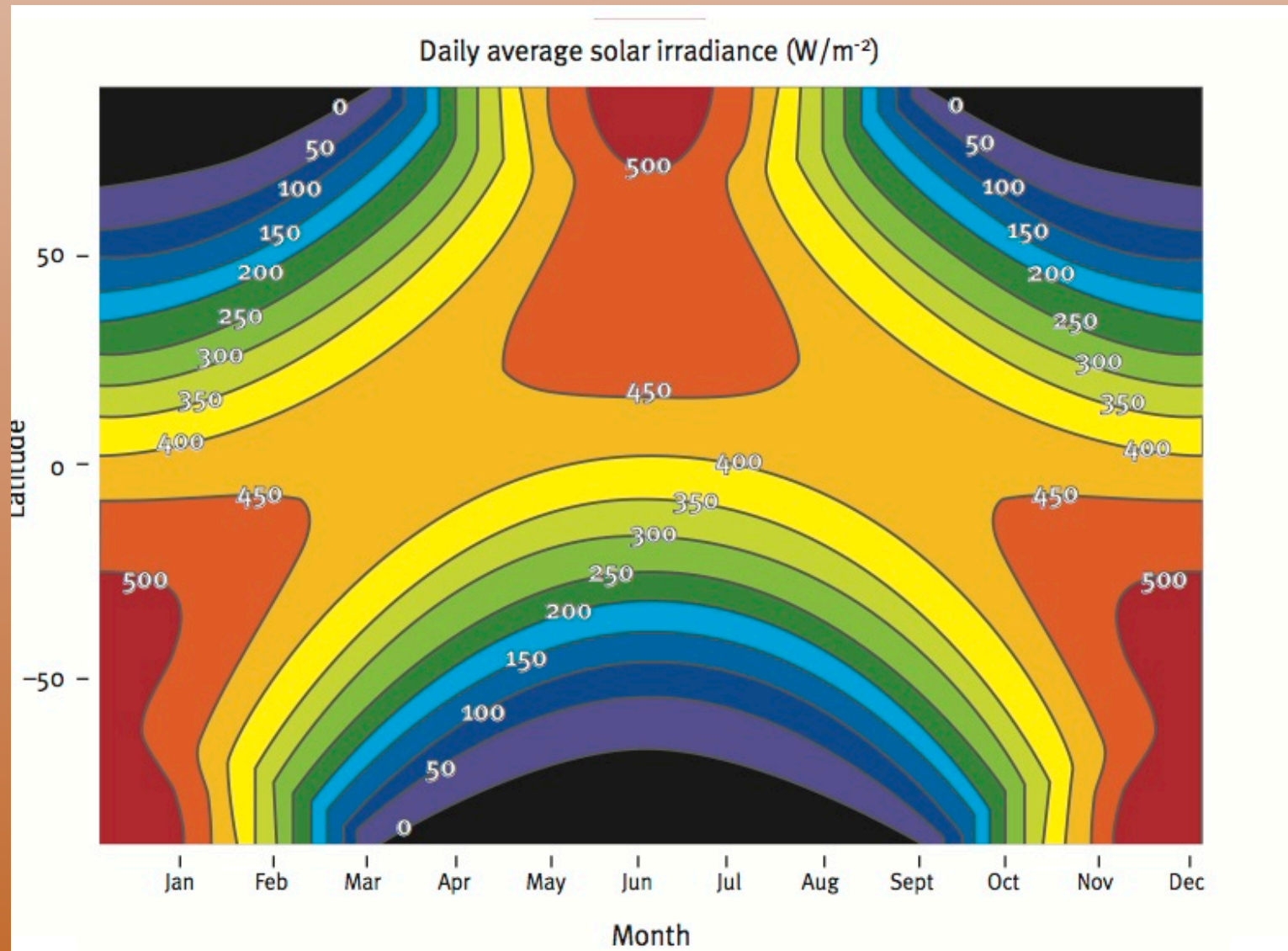
$$Q_0 = S(d_m / d)^2 \int_{\text{sunrise}}^{\text{sunset}} \cos(Z) dt \quad \text{Total daily insolation}$$

$$Q_0 = \frac{24}{\pi} S(d_m / d)^2 \left( \int_0^H \sin \phi \sin \delta dh + \int_0^H \cos \phi \cos \delta dh \right) =$$

$$\frac{24}{\pi} S(d_m / d)^2 (H \sin \phi \sin \delta + \cos \phi \cos \delta \sinh)$$

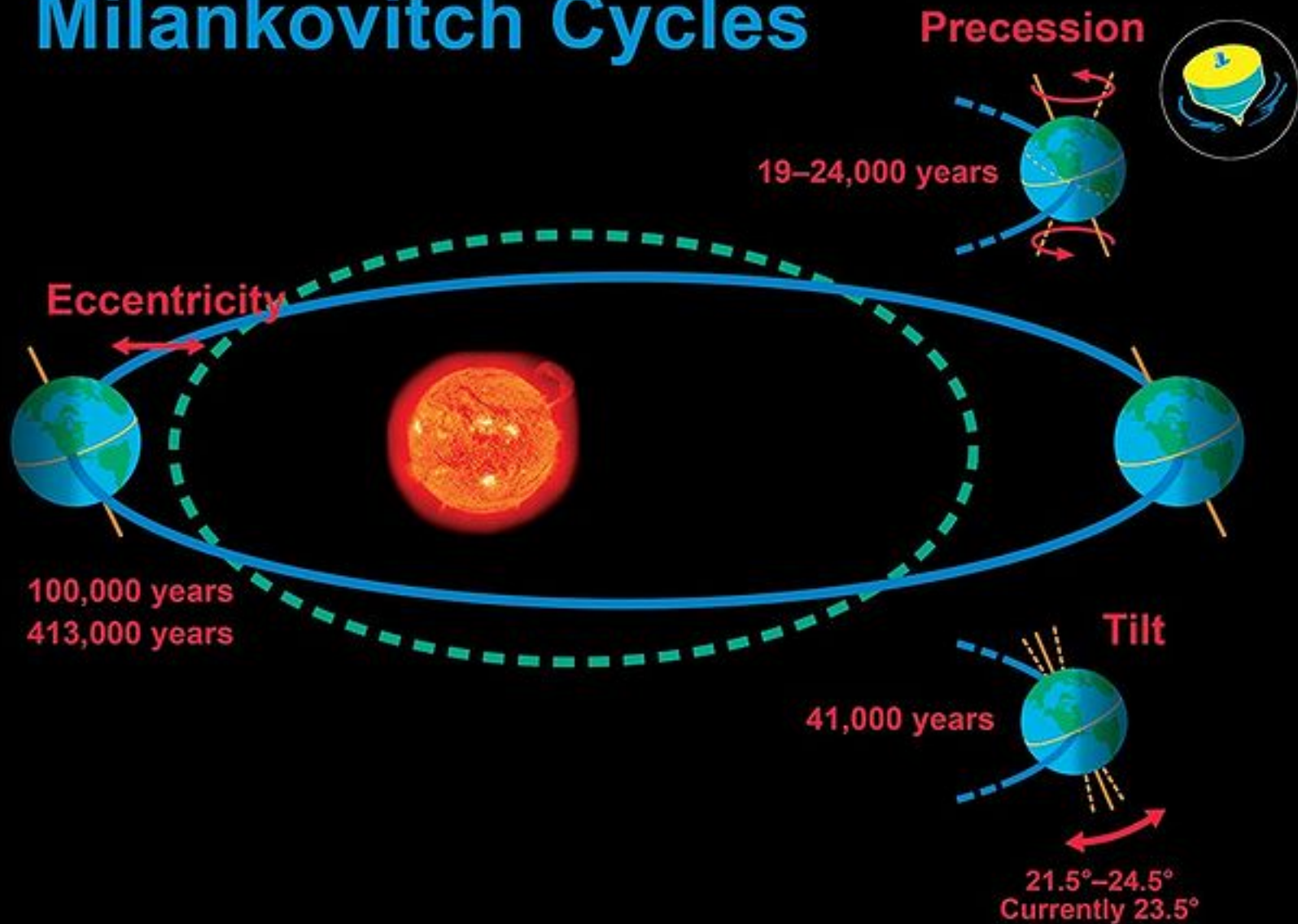
H=rotation\_rate \*(local\_time-12)

# Distribution of Solar radiation at TOA

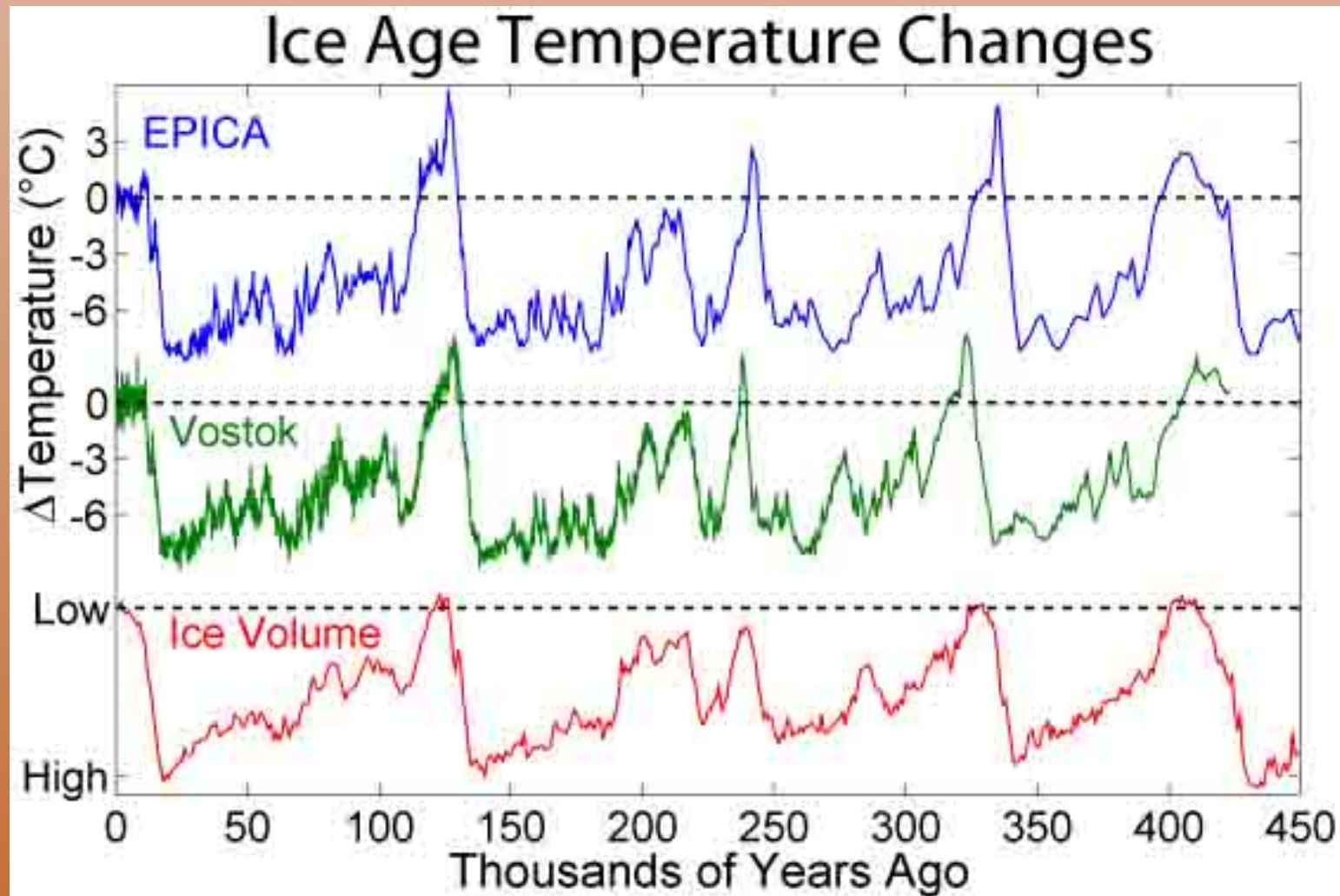


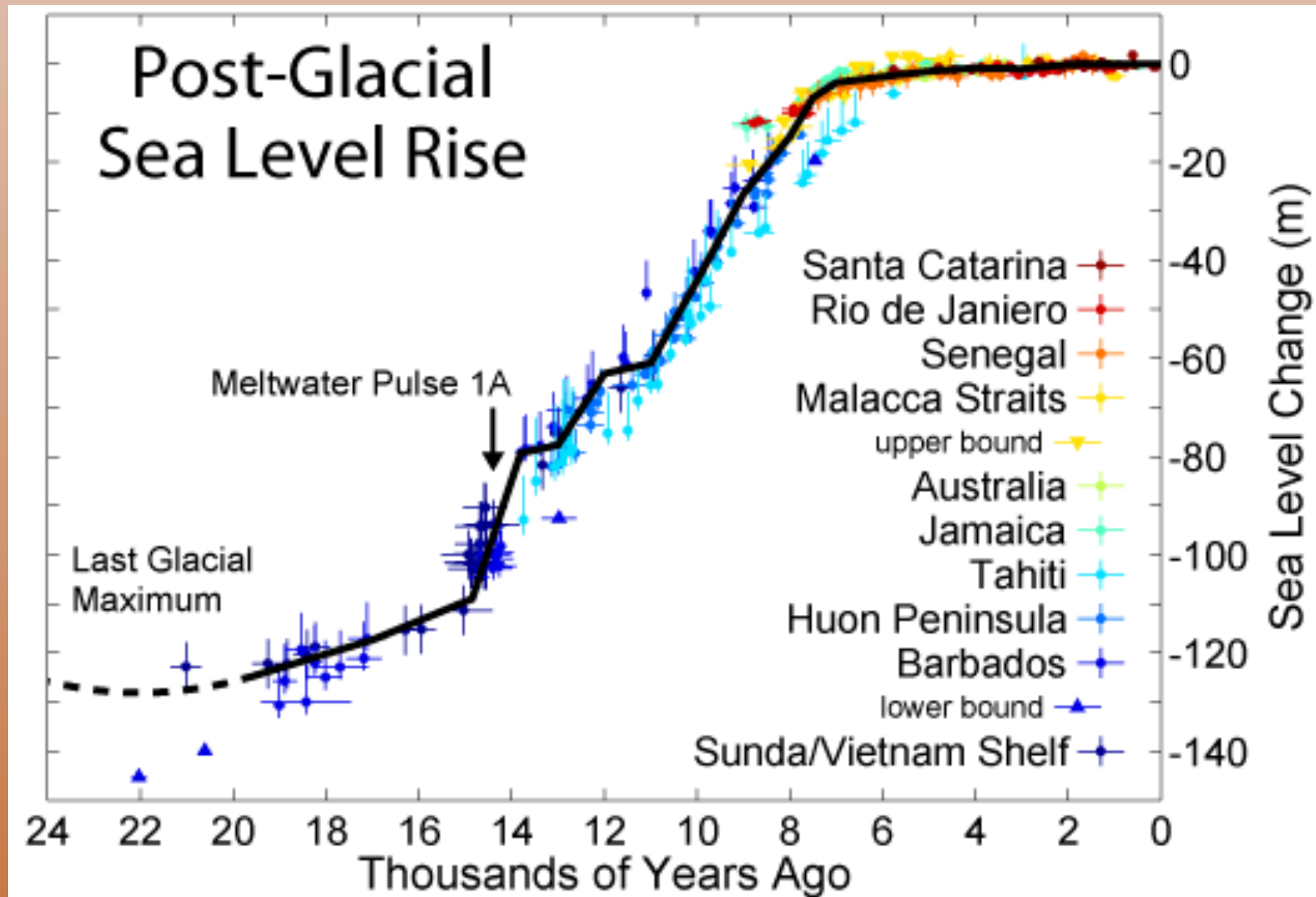


# Milankovitch Cycles



# Paleoclimate proxies



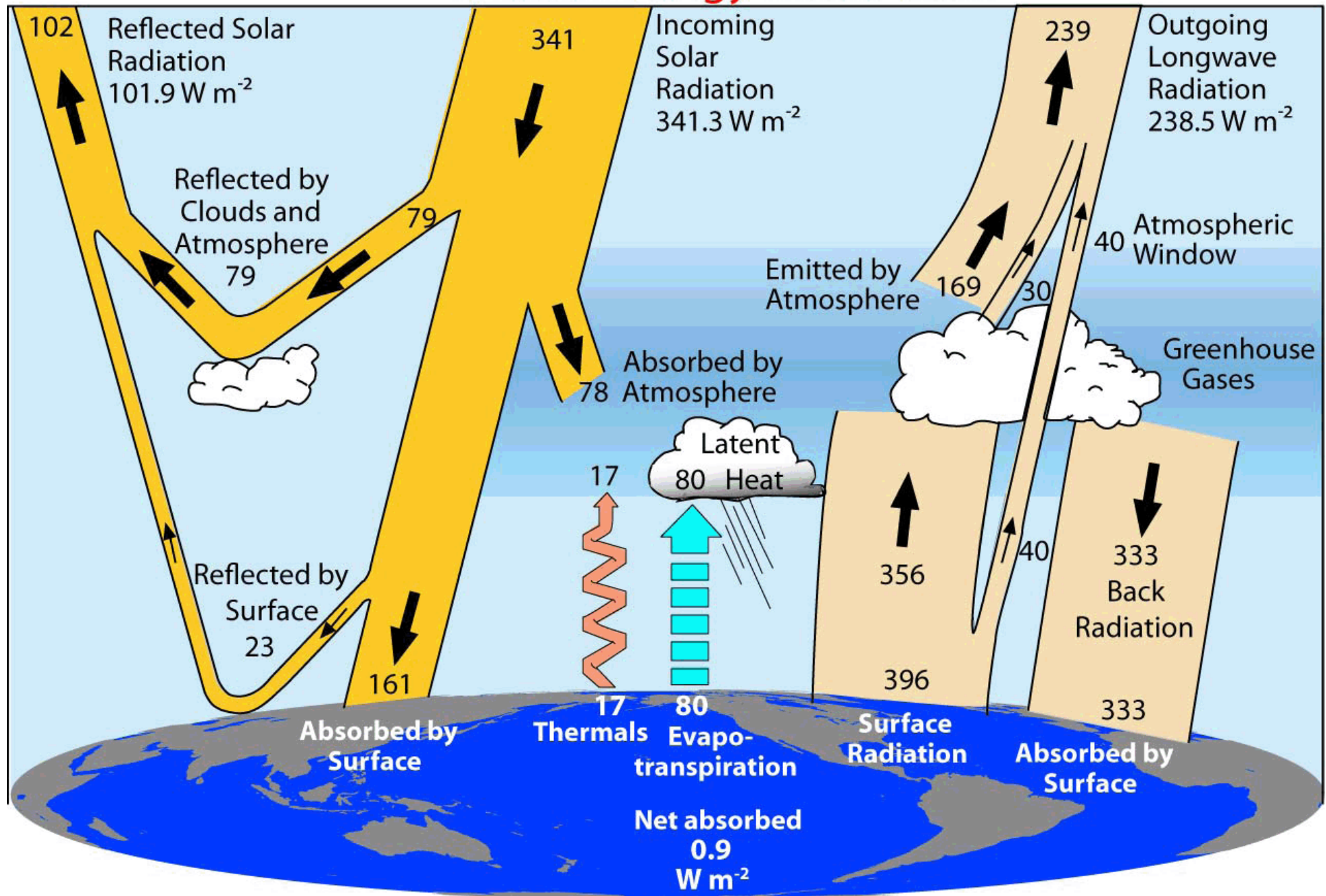


10 W/m\*\*2 at the pole( $\Delta S$ )

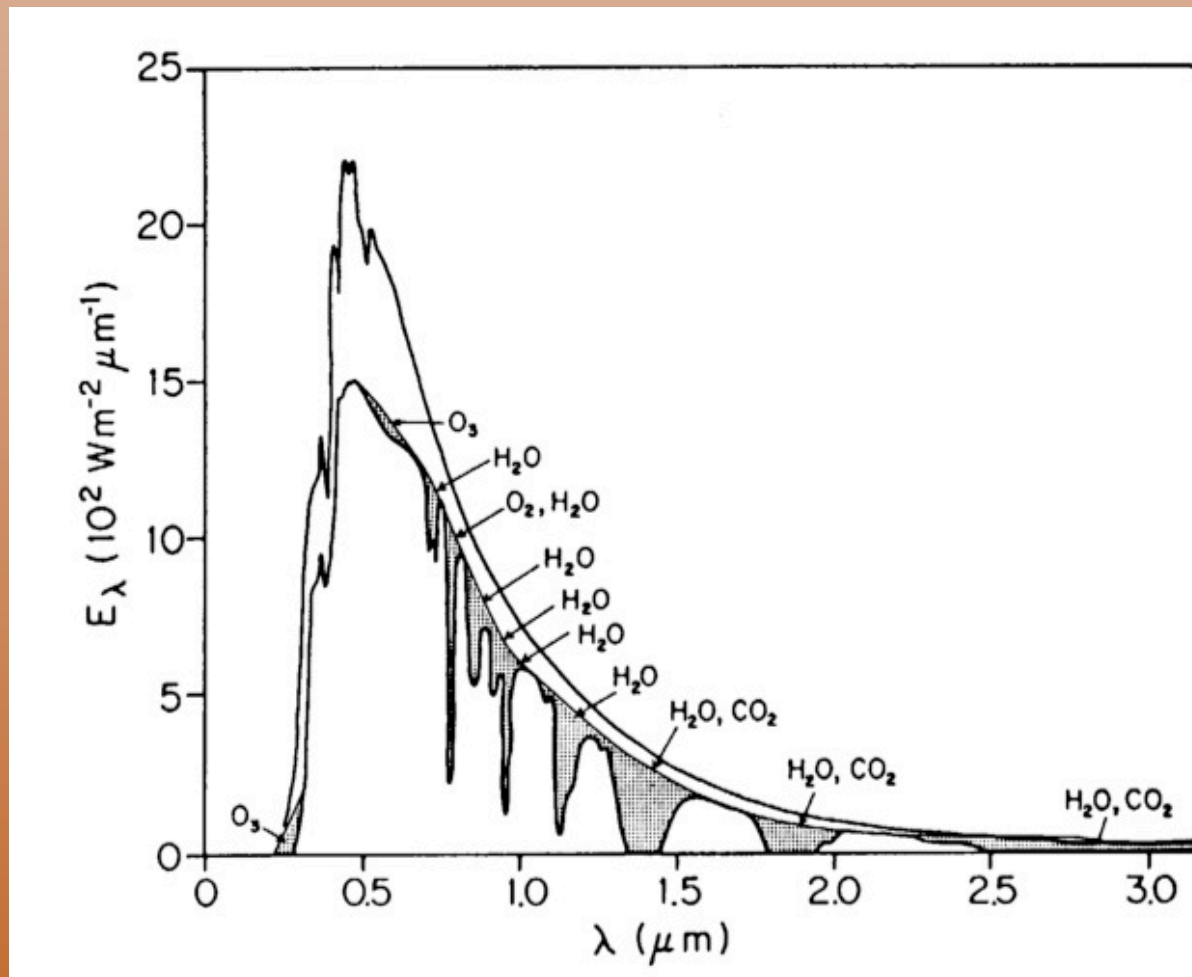
5 deg (global  $\Delta T$ )

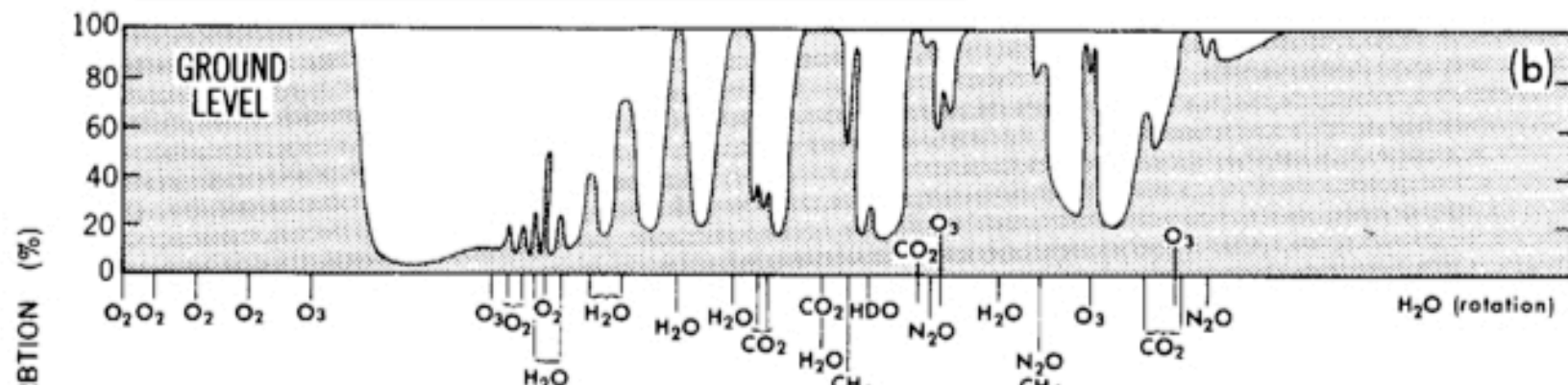
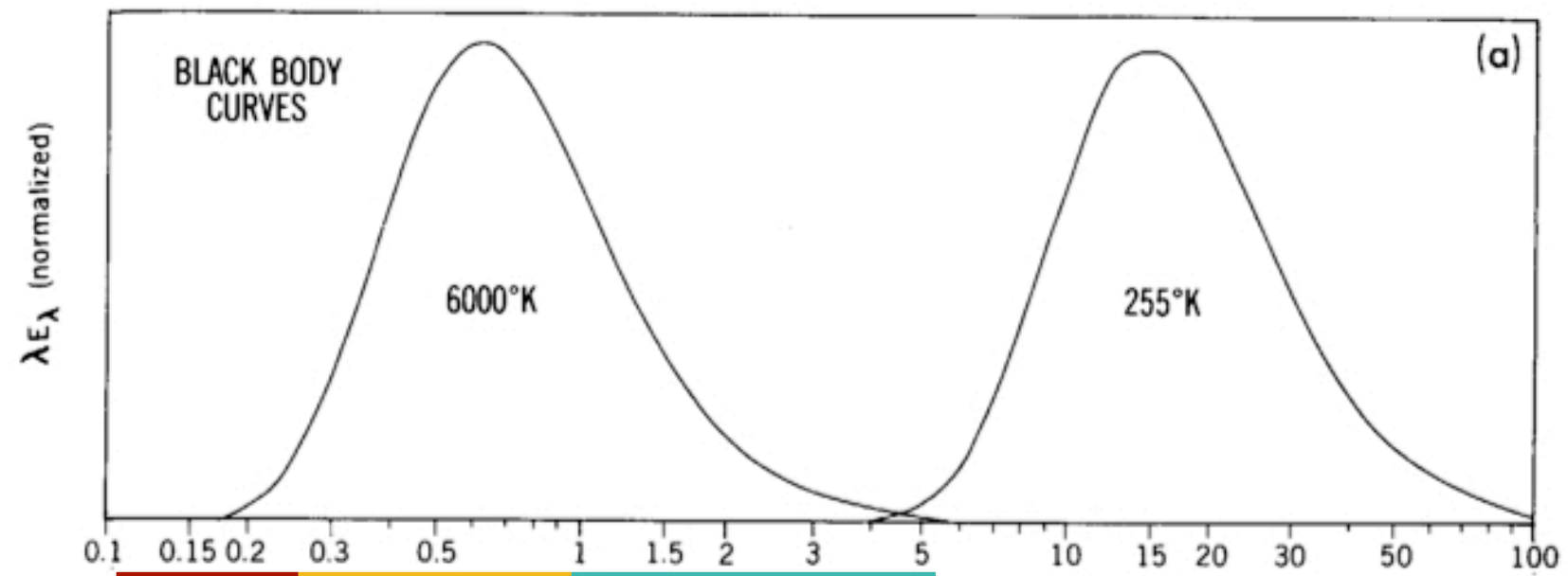


# Global Energy Flows $\text{W m}^{-2}$ TOA Balance



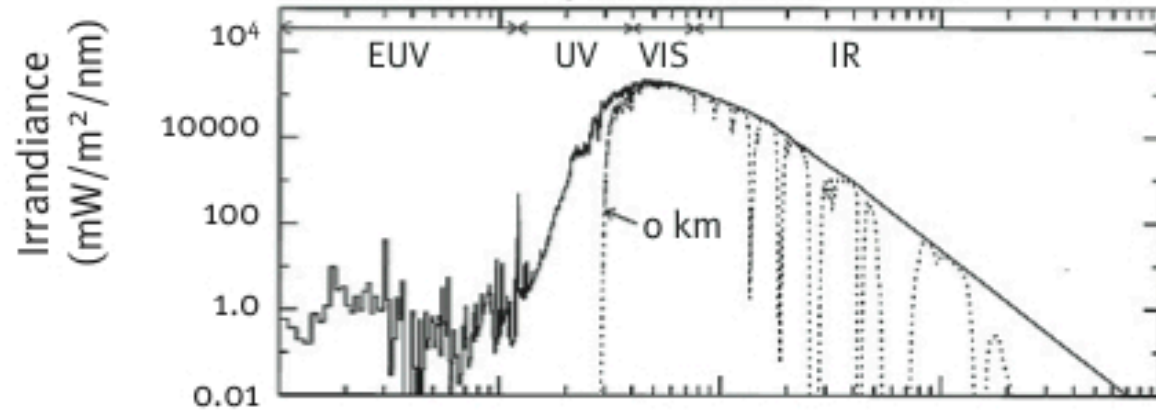
# Spectral distribution of Solar Irradiance at TOA and sea level



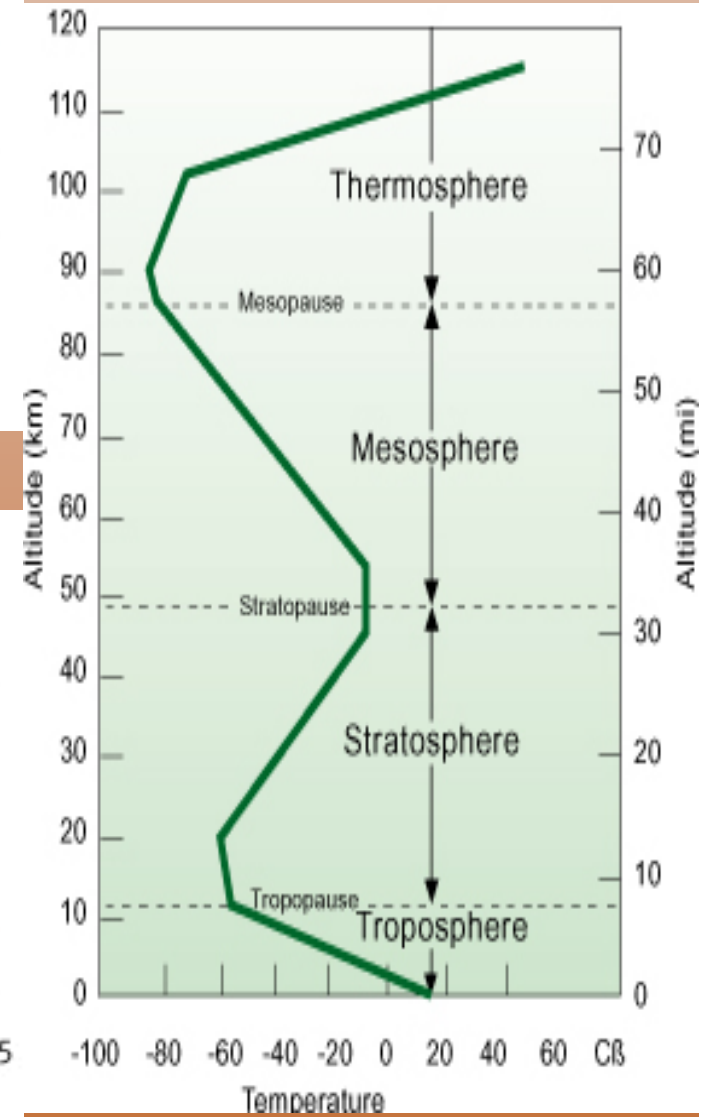
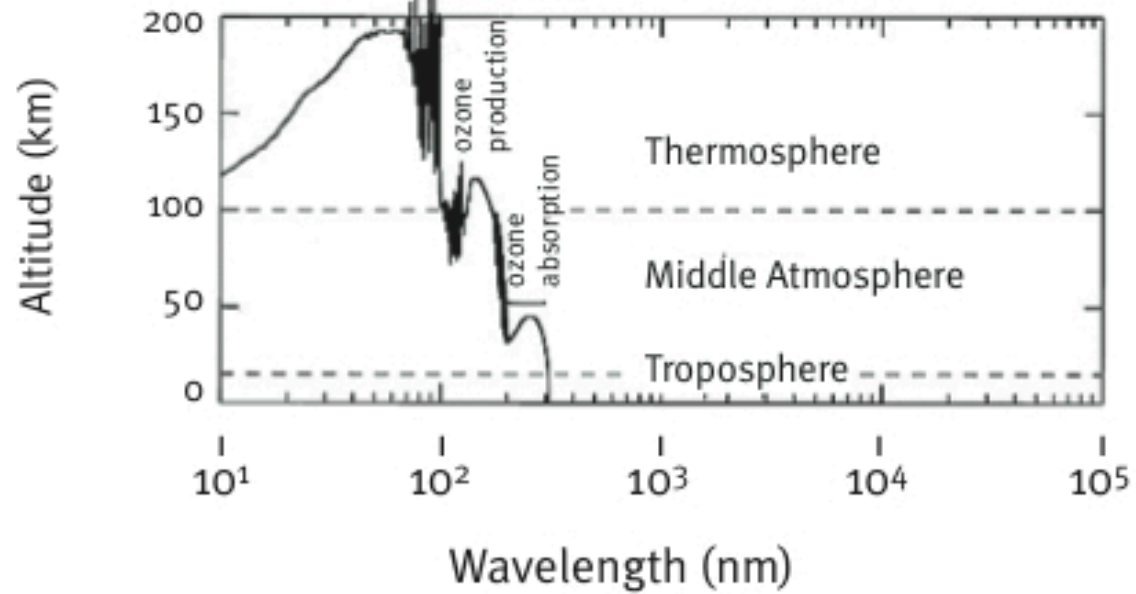


micron

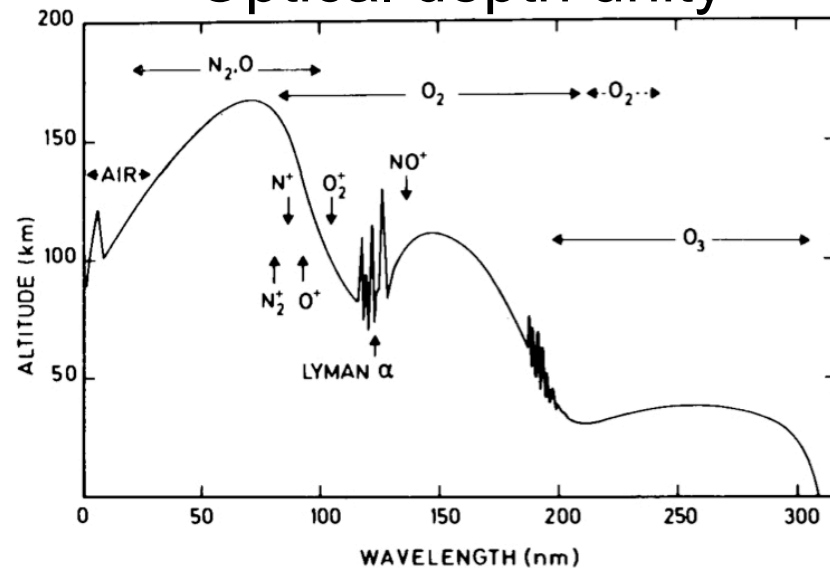
(a) Solar spectral irradiance



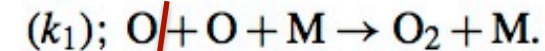
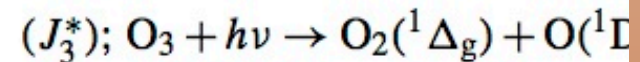
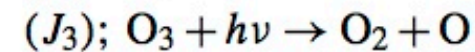
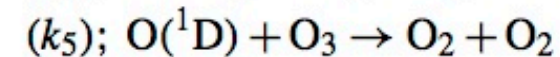
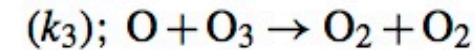
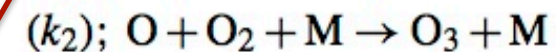
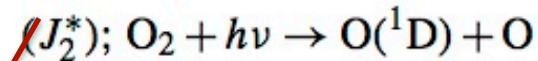
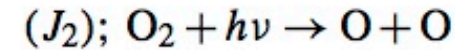
(c) Terrestrial optical depth unity



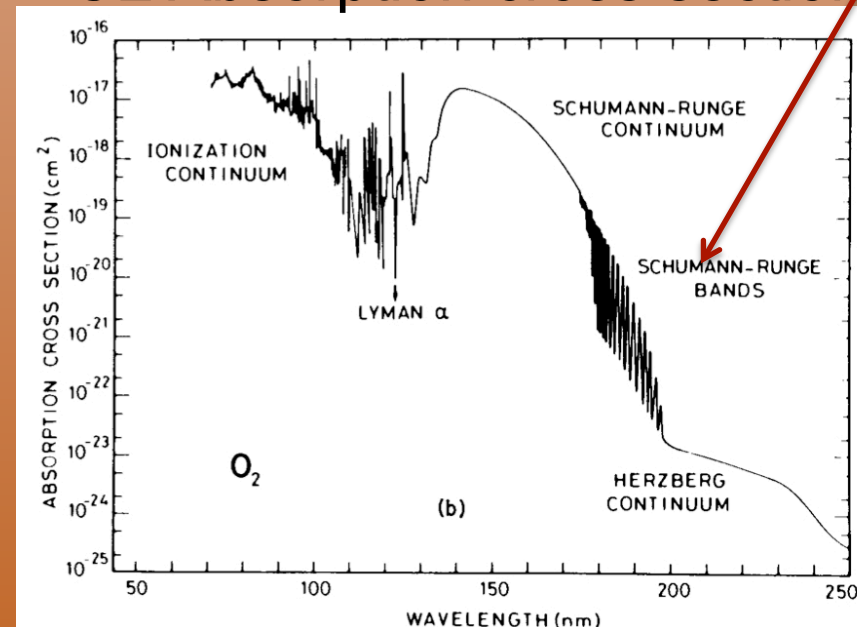
## Optical depth unity



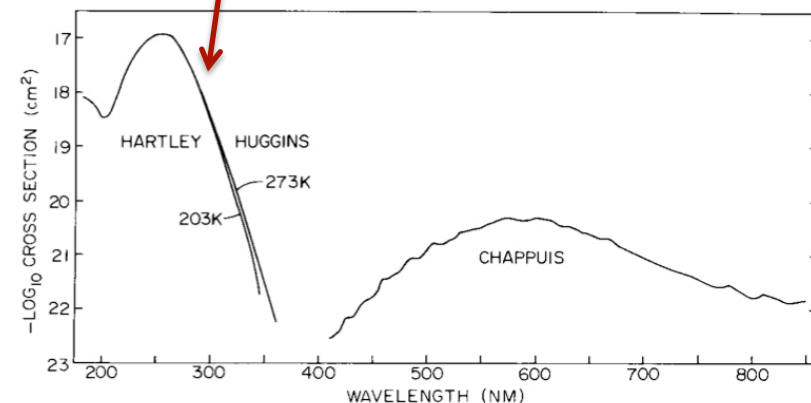
## Chapman Cycle



## O2 Absorption cross section

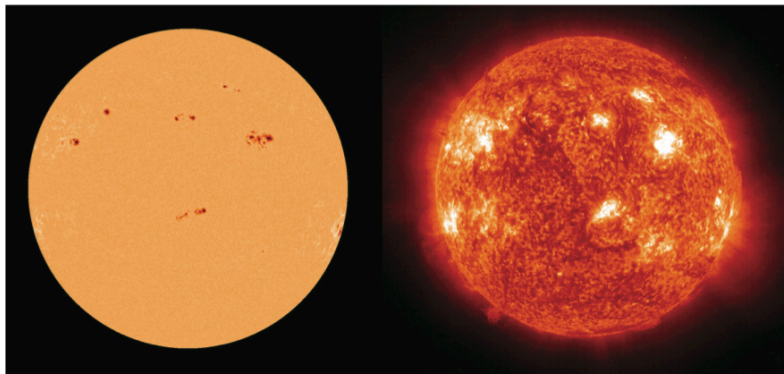
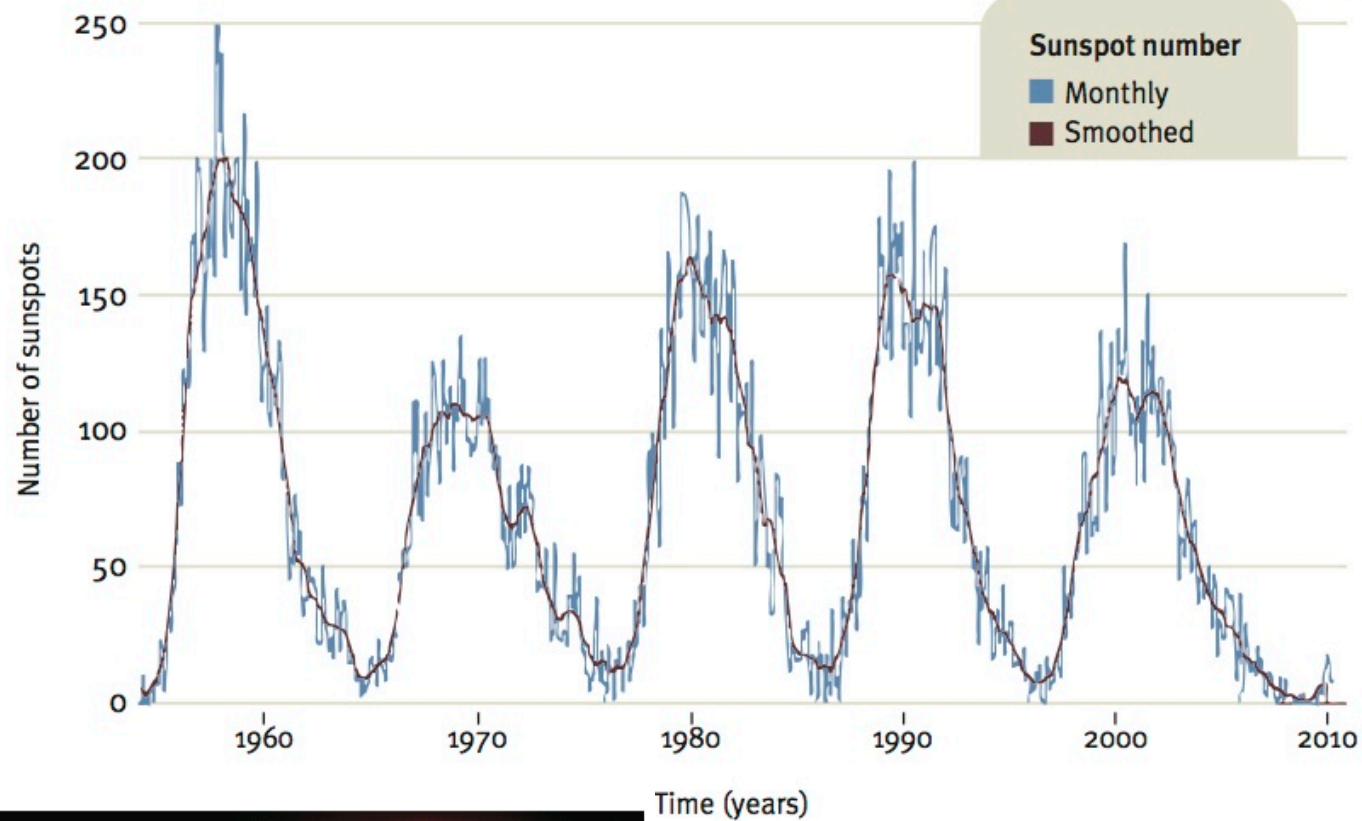


## O3 Absorption cross section





# Indicators of solar activity : sunspot numbers



**Figure 8.** The Sun observed on 20 March 2000, near the maximum of the most recent solar activity cycle. Left: in visible light, showing sunspots; Right: in the ultraviolet, revealing areas of intense activity associated with the sunspots<sup>10</sup>.

# Indicators of solar activity

Sunspot

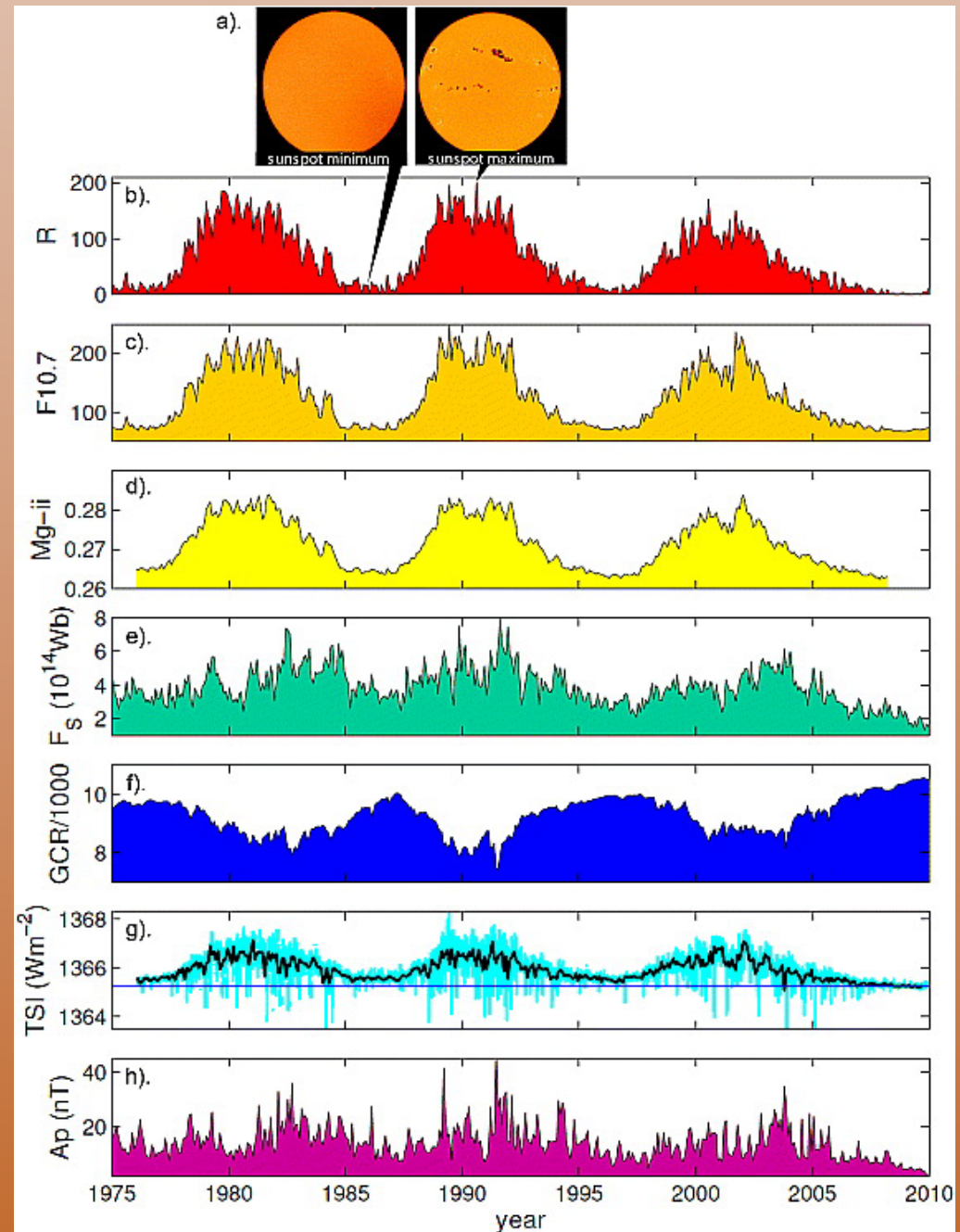
10.7 cm radio flux

Mg II

GPC

TSI

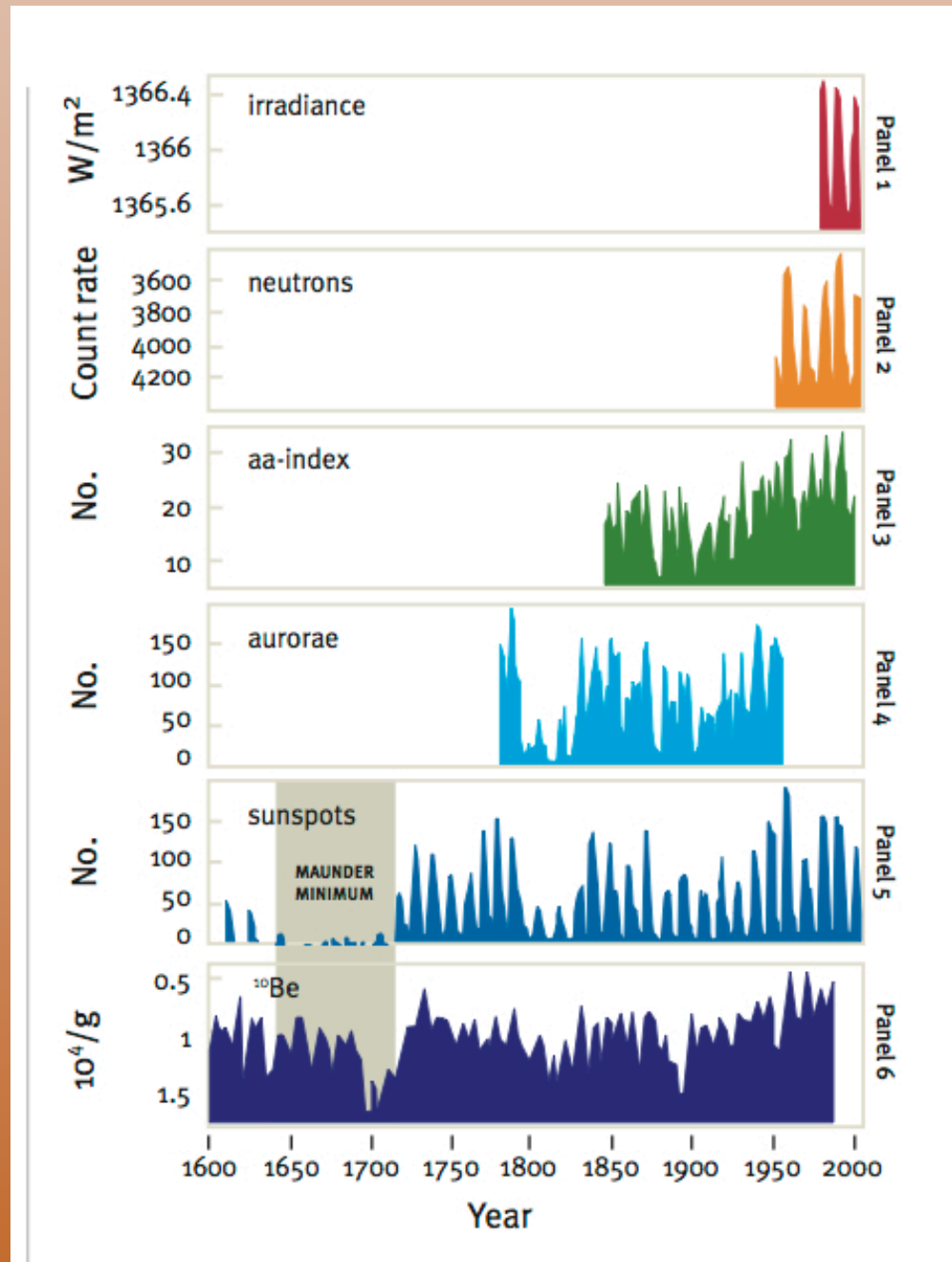
Index of gm disturbance



# Indicators of solar activity

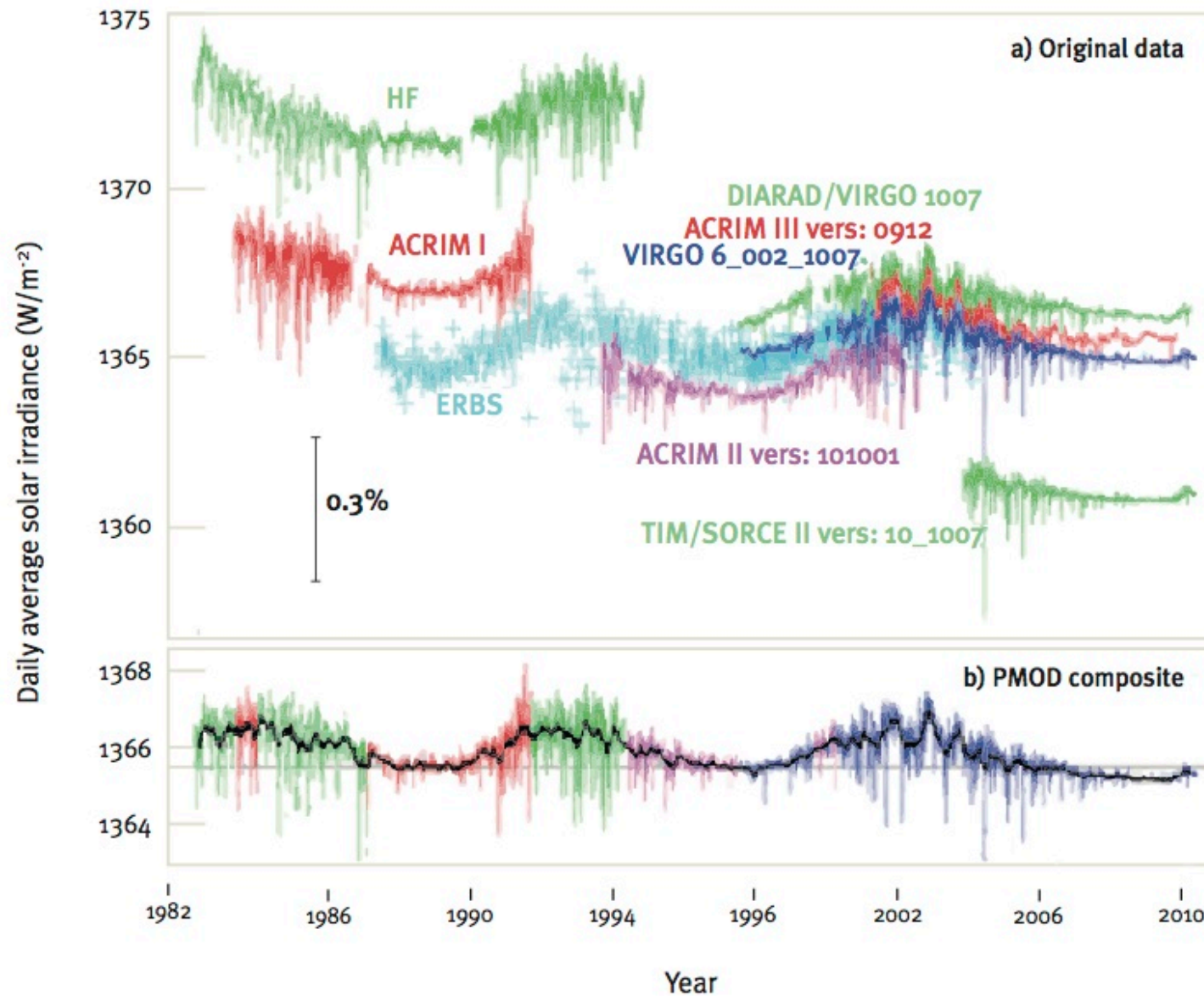
TSI

Sunspot numbers



- Proxies for solar irradiance are used for climate-sun relationship
- If we could get understanding of how sun's magnetic activity is related to solar irradiance, we could reconstruct past solar irradiance
- Predict solar irradiance → effects on climate

# The first candidate: TSI

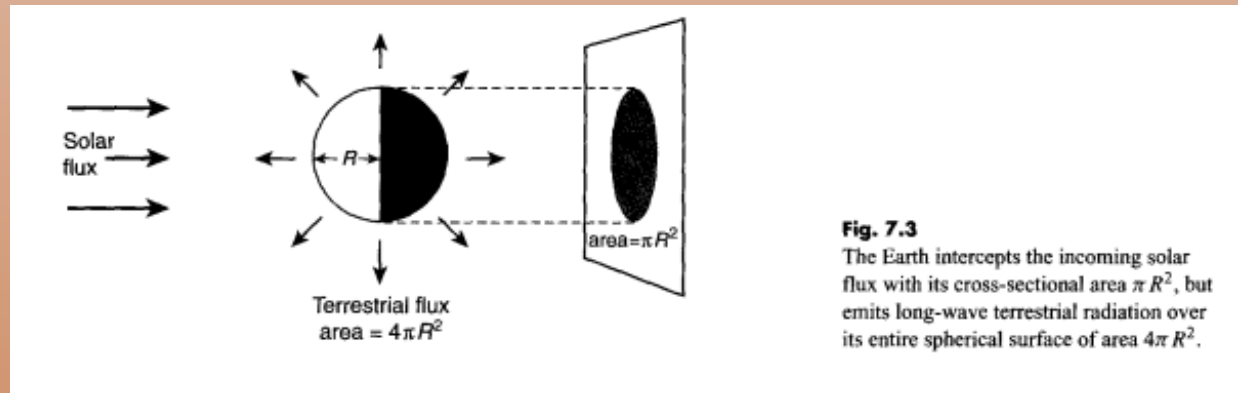


**Figure 10.** (a) Daily-averaged total solar irradiance ( $\text{Wm}^{-2}$ ) from 1978 to present: all measurements made from satellites. (b) Composite record obtained by inter-calibration of the data from the individual instruments<sup>16</sup>.



Solar energy absorbed by the Earth =  $(1 - A)S/4$ ,  
where  $A$  is the Earth's albedo and  $S$  is the total solar irradiance

$S_{\text{max}} - S_{\text{min}}$  = of about  $1 \text{ W m}^{-2}$



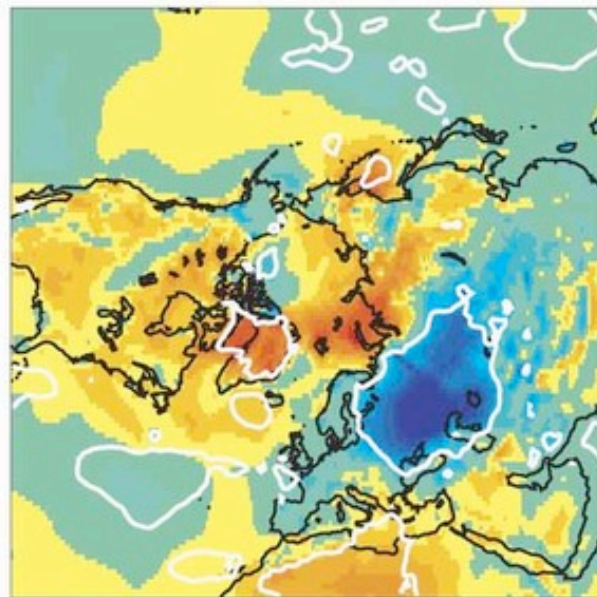
$S = 1366 \text{ W m}^{-2}$  and  $A = 0.3$ ,  
the solar power available to the Earth system =  $239 \text{ W m}^{-2}$   
with an 11 year SC variation of  $\sim 0.17 \text{ W m}^{-2}$ , or  $\sim 0.07\%$

Impact at the surface :  $DT_s = I \cdot DF$ ,  $I = \text{climate sensitivity } 0.5 \text{ K (W/m}^2\text{)}^{-1} \rightarrow DT_s = 0.07 \text{ K}$

What do Observations show ?



-3 -2 -1 0 1 2 3  
Reanalysis sea-level pressure difference (hPa)



-2.0 -1.5 -1.0 -0.5 0 0.5 1.0 1.5 2.0  
Reanalysis temperature difference (K)

Sea Level Pressure

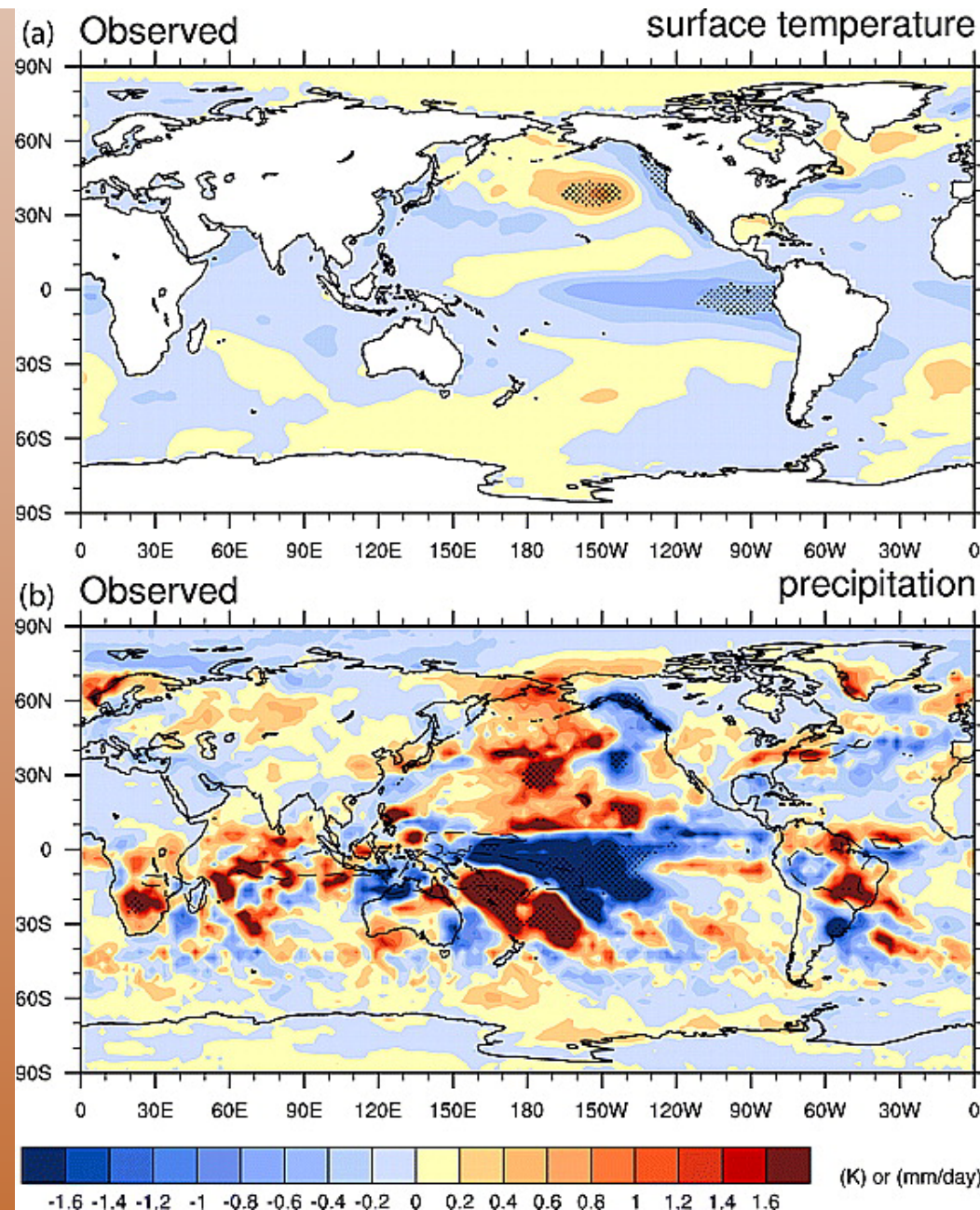
EXTRATROPICS  
SLP and temperature  
solar cycle signal at the  
surface from reanalyses

Surface temperature

Ineson et al., Nature Geoscience 2011

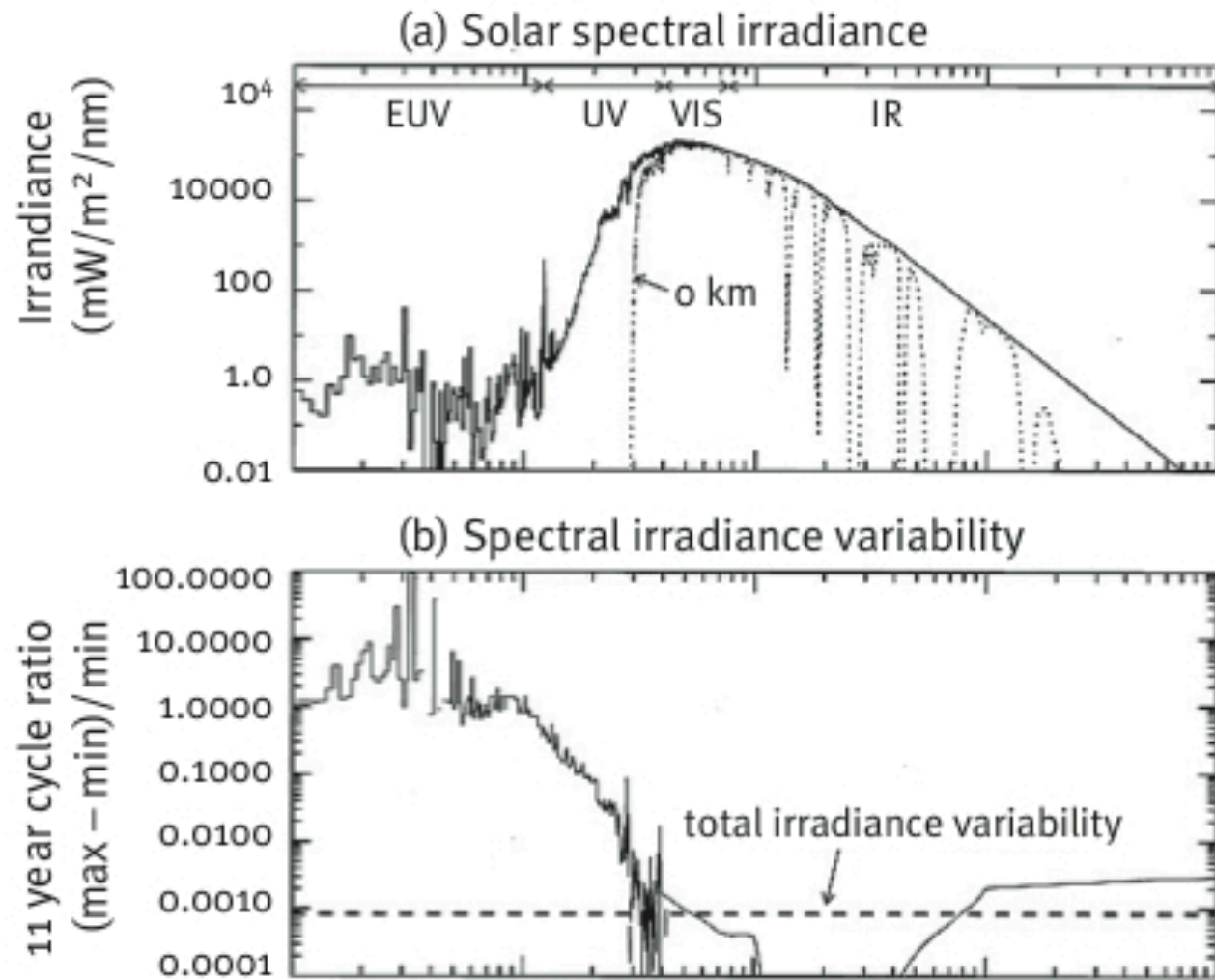
# TROPICS

Meehl et al., Science



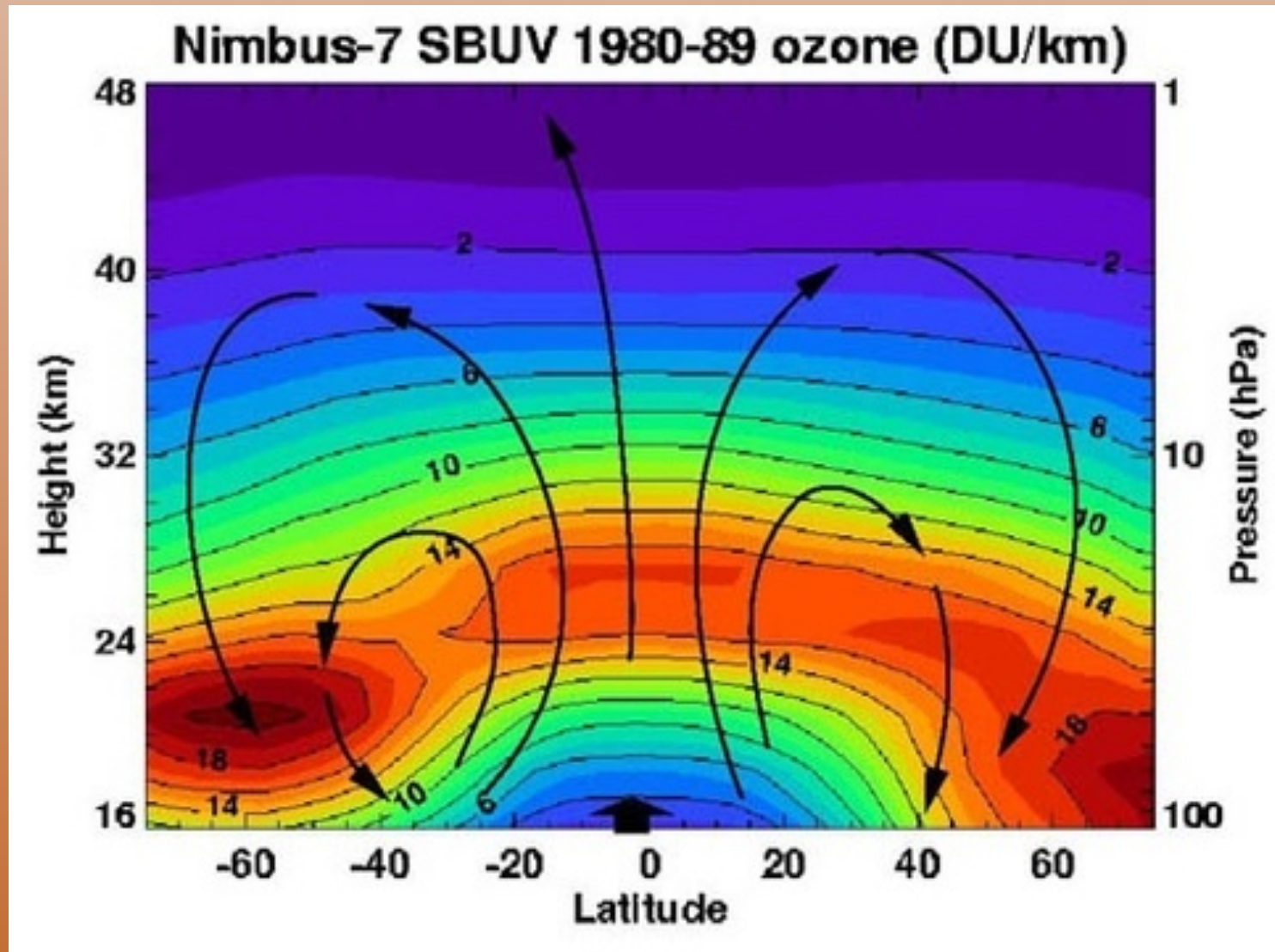
Regional responses larger than expected global mean :  
Feedbacks ? Amplifying mechanism ??

## The second candidate: SSI



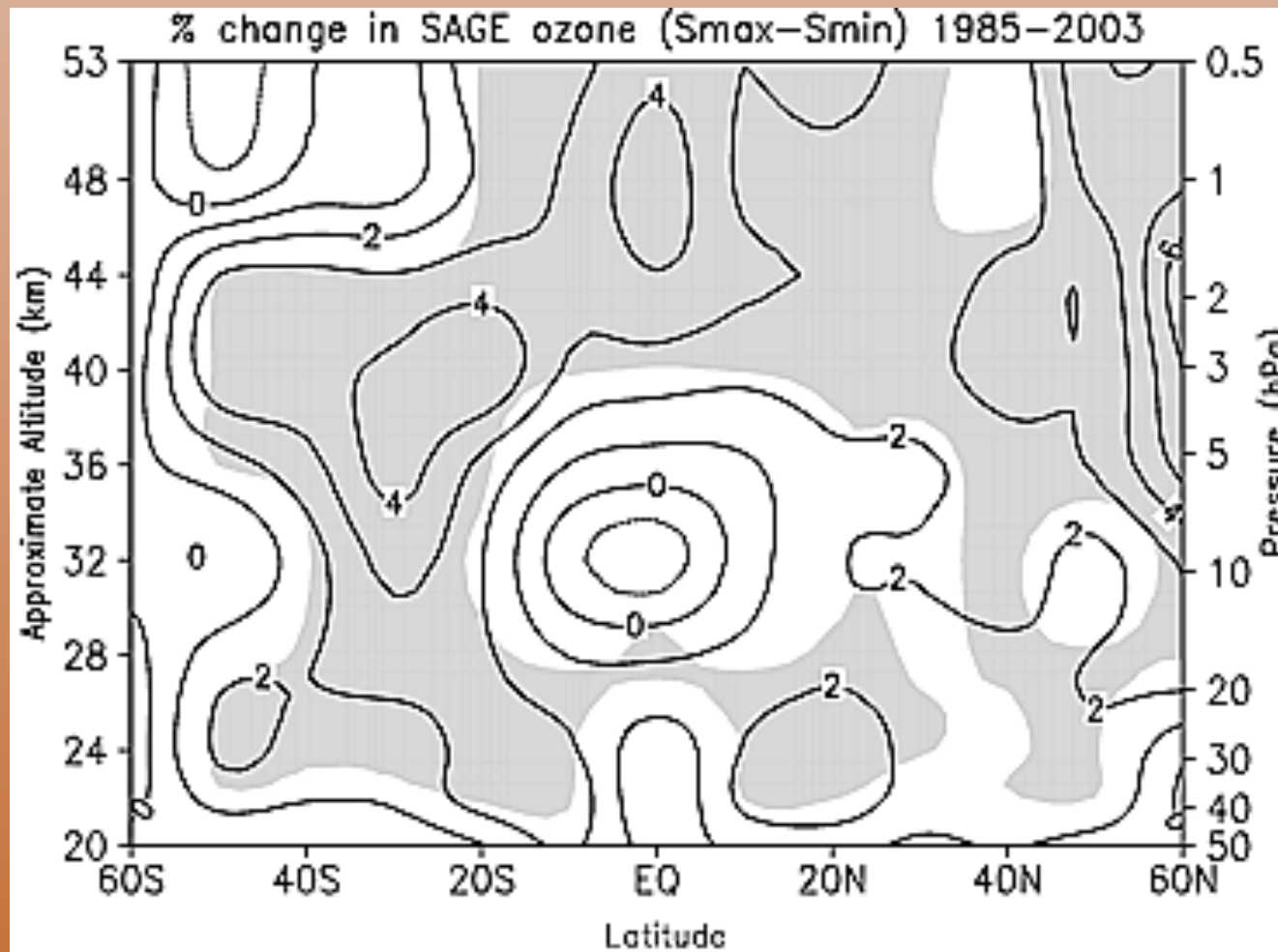


# Stratospheric Ozone Climatology- year mean



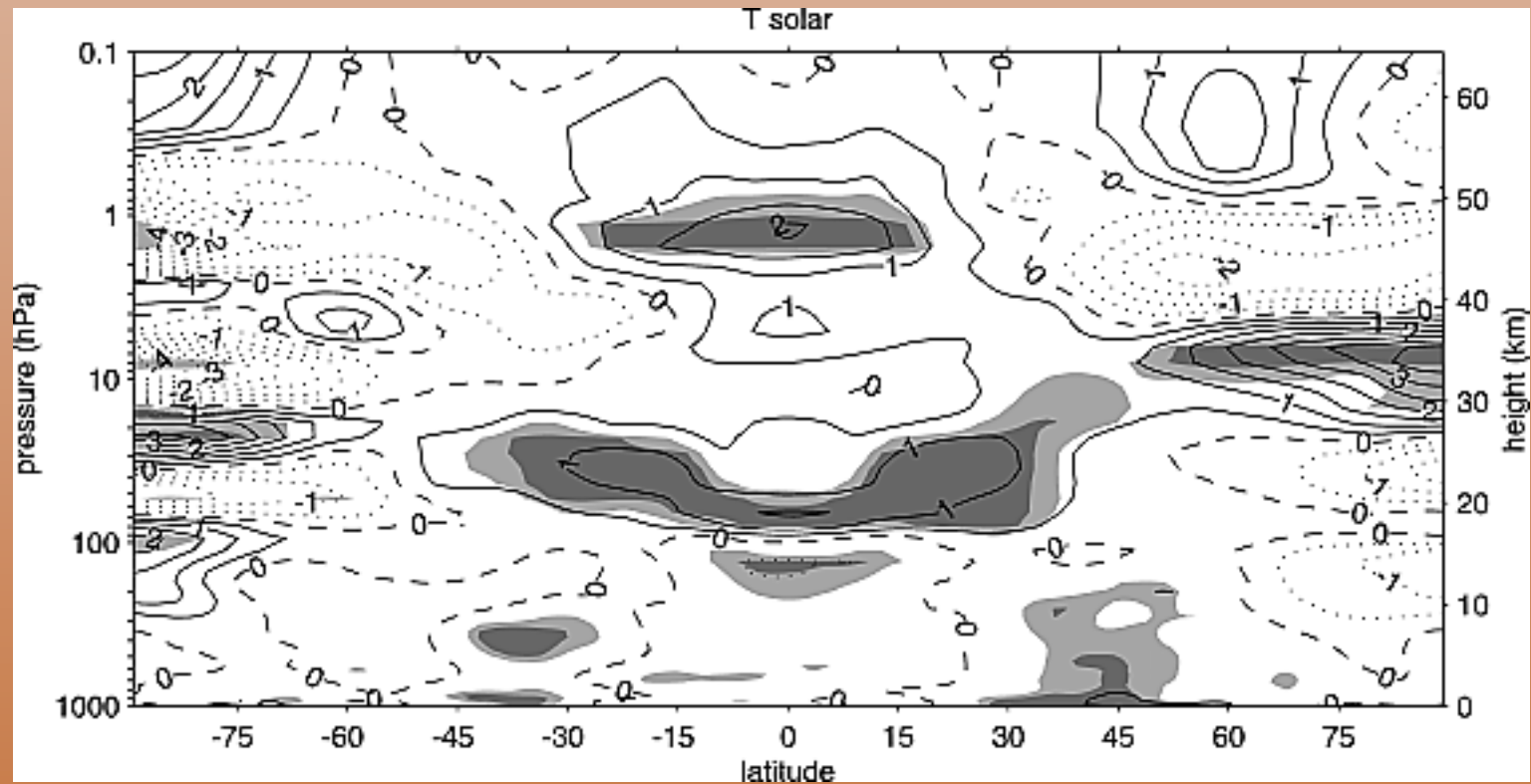


## 11-yr Solar signal on Ozone



Soukharev and Hood 2006

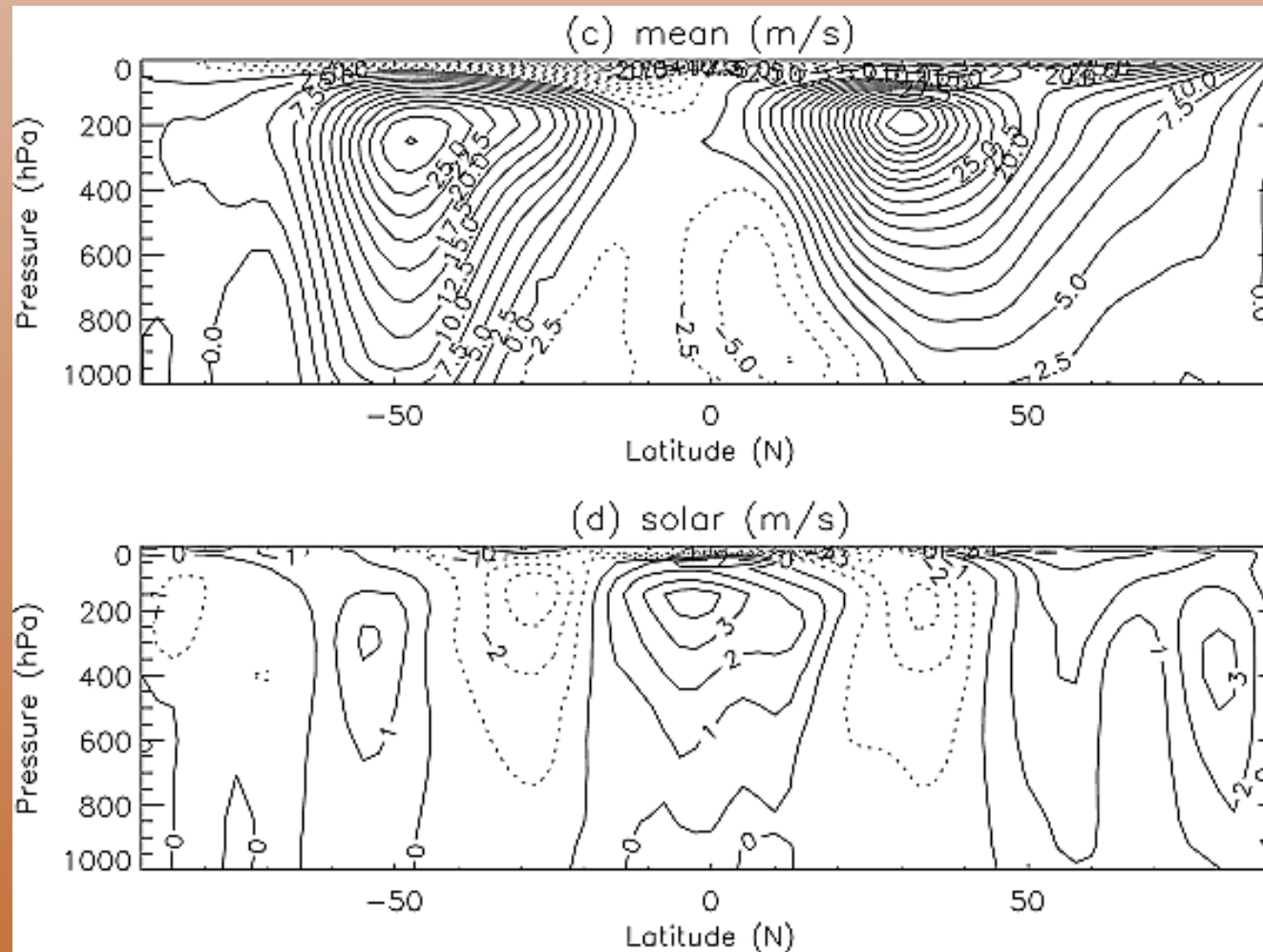
# 11-yr Solar signal on Temperature



Change the temperature → change in the circulation

Gray et al., 2010

## 11-yr Solar signal on winds



Change the temperature → change in the circulation

## **The possible Mechanisms**

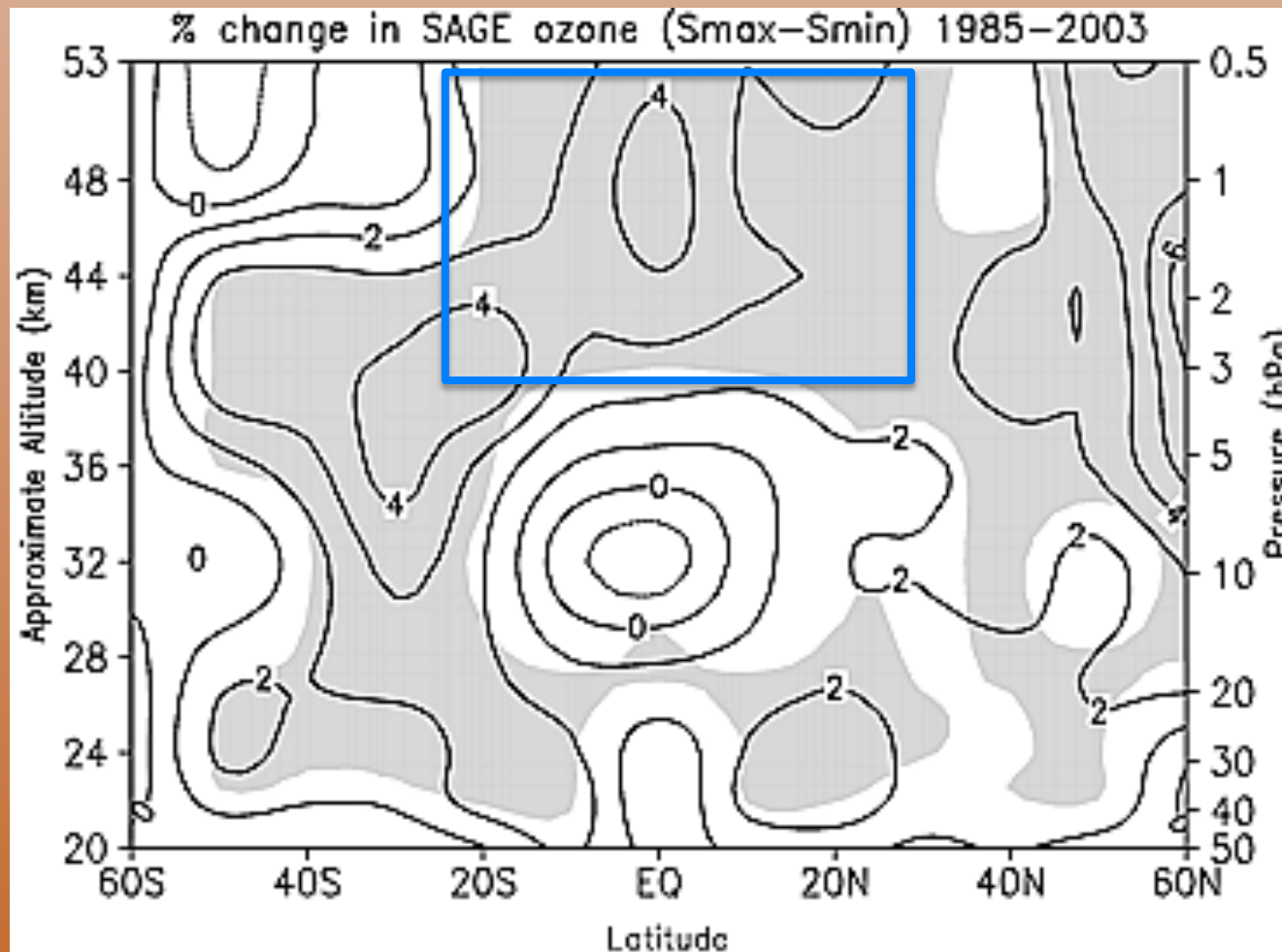
Top-down mechanism, 1. through the stratospheric ozone – polar route in winter, 2. at equator, all year (SSI)

Bottom-up mechanism, in the tropics (TSI)

Interaction between GCRs and clouds

## Top-Down

+ phase Solar cycle in the UV → + Ozone Tropics, middle atm

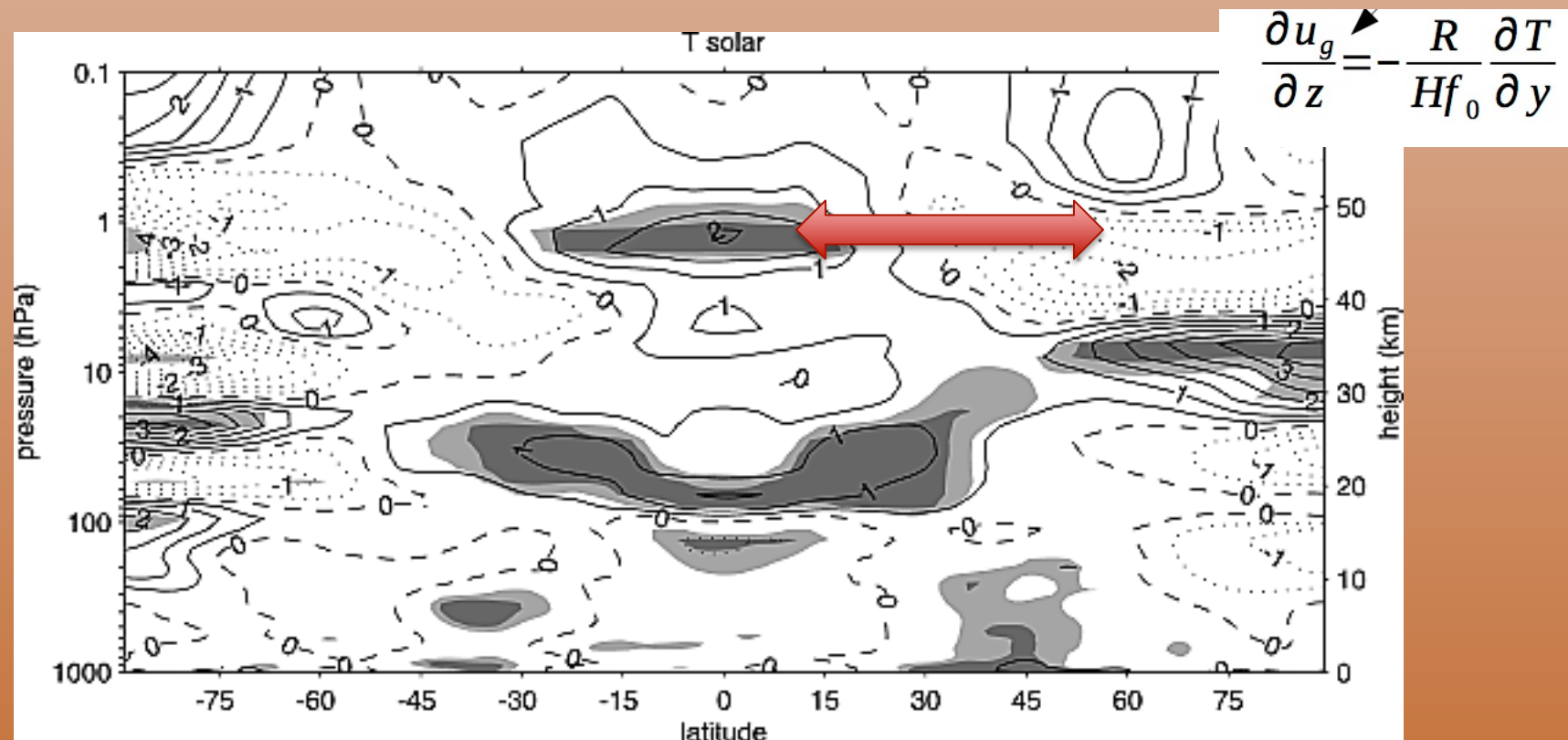




## Top-Down, polar route

+ phase Solar cycle in the UV → + Ozone Tropics, middle atm

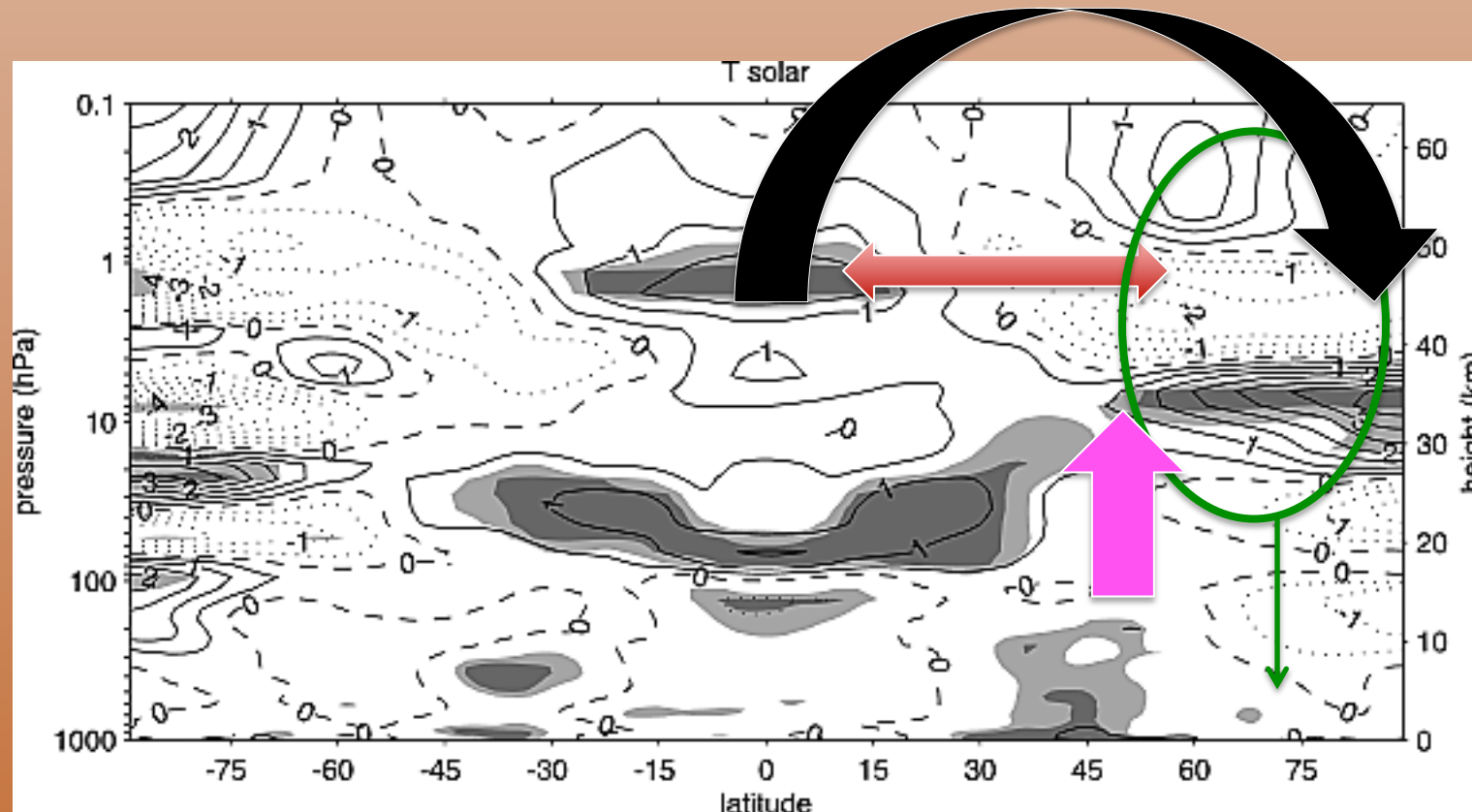
→ Warming of tropical lats → impact on the meridional T gradient



## Top-Down, polar route

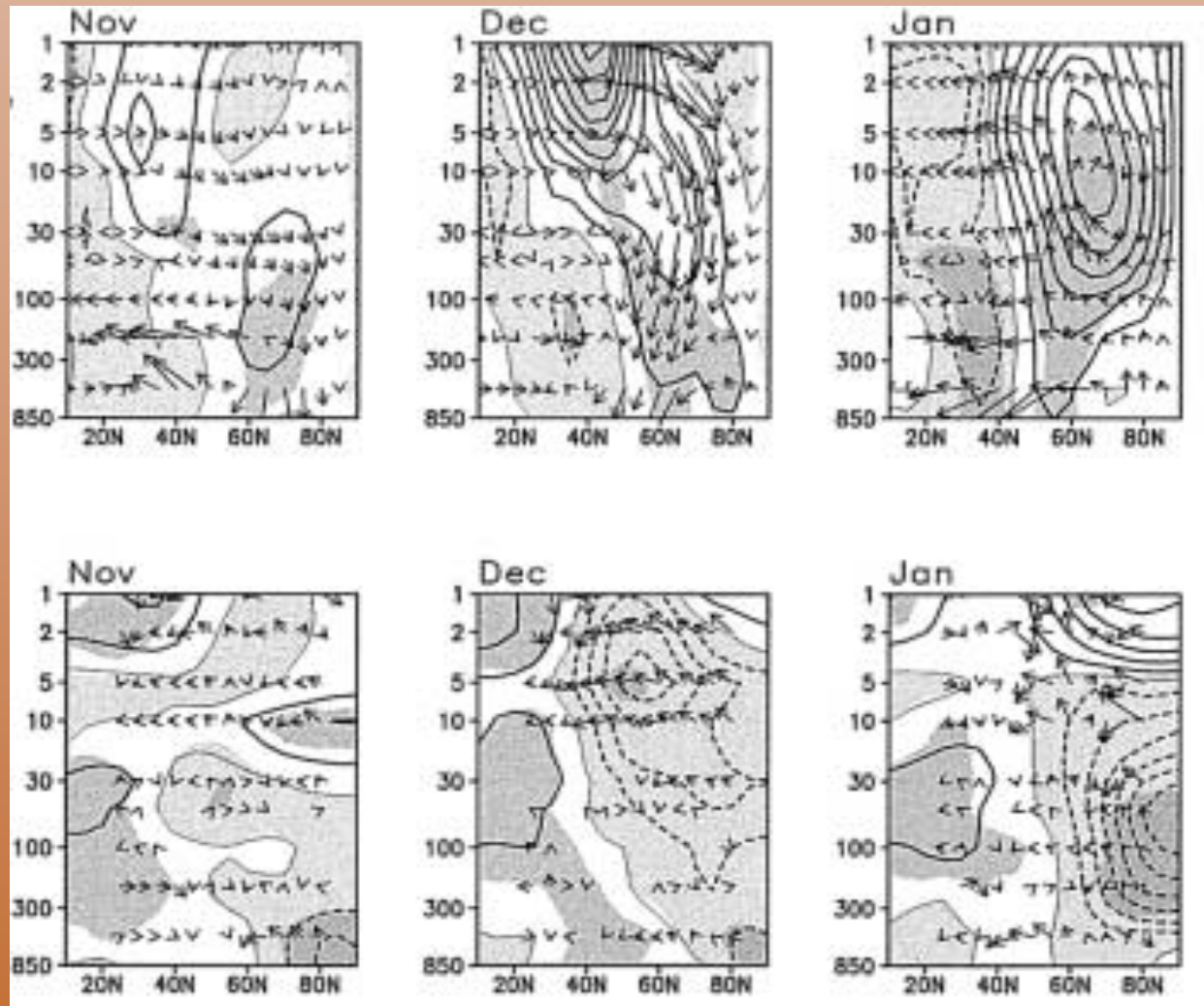
+ phase Solar cycle in the UV → + Ozone Tropics, middle atm

→ Warming of tropical lats → impact on the meridional T gradient



→ Impact on the winds (+) → Planetary wave propagation →  
→ Meridional circulation(-), wind anomalies propagate down

## Evolution of wind and T, beginning of winter



Wind

Temperature

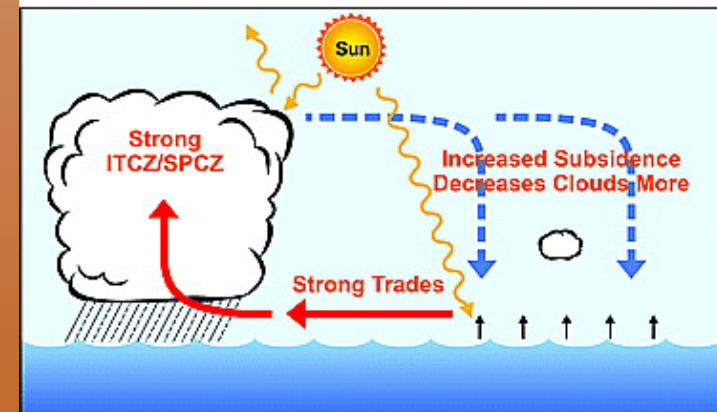
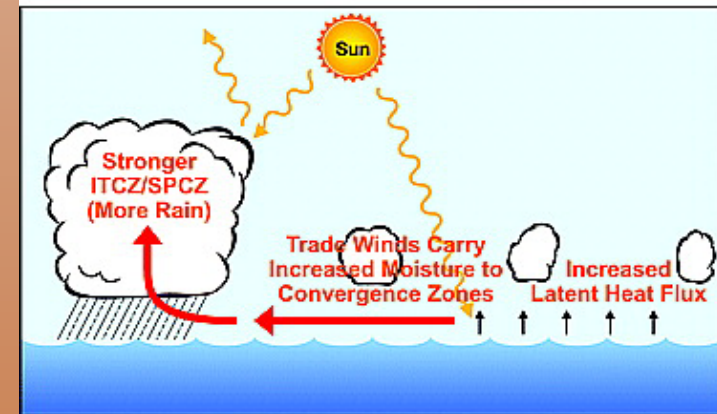
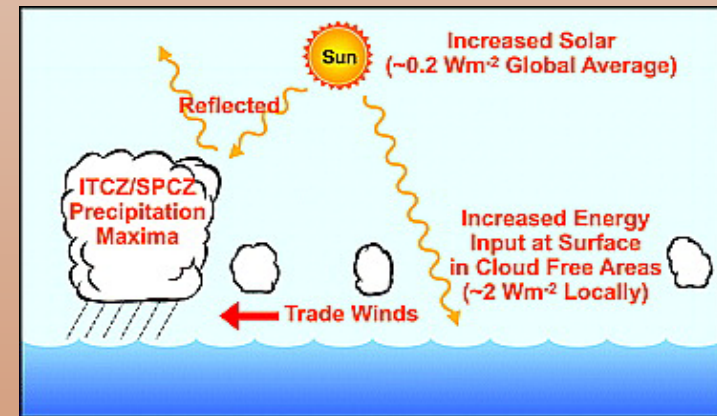
Kodera and Kuroda, 2002

## Bottom-up:

Absorption of solar radiation over cloud free subtropical ocean, larger during Smax →

larger evaporation and more moisture transported into precipitation zones → larger upward vertical motion

→ stronger trade winds → ocean upwelling → colder SSTs → feedback reducing clouds



## **Charged particles:**

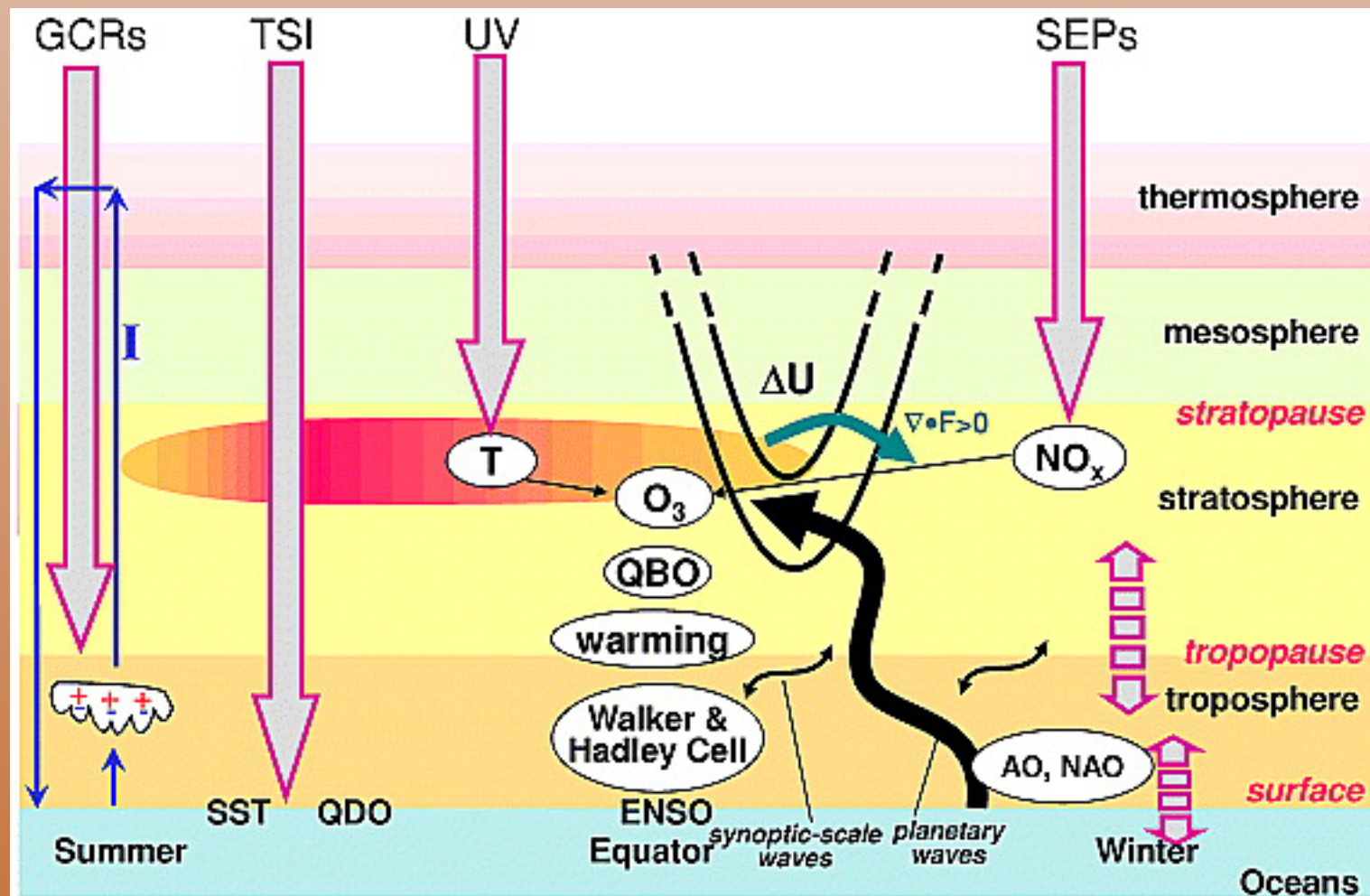
Cosmic rays generate ions. Two different routes might result in this ionisation influencing cloud cover:

- preferential growth of ionised particles to a size which is energetically favourable for cloud droplet formation, then other processes are needed to really have enhanced CCN
- ionisation might influence cloud cover through changes in the global electrical circuit where currents flowing between the surface and the ionosphere initiated by thunderstorm activity. Near the edges of clouds charge can accumulate and this can influence both evaporation of cloud droplets and interactions between them

Plausible mechanism, but no evidence

EPP influence chemical constituents (HOx, NOx..) that is transported in the polar vortex may impact polar O<sub>3</sub>

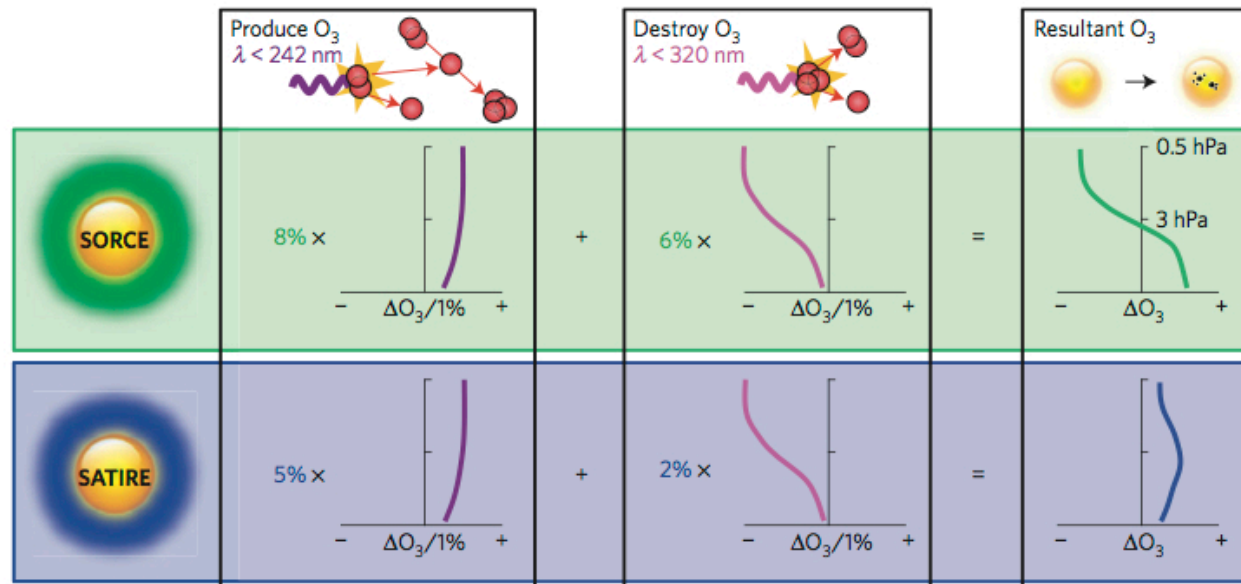




# Uncertainty in the SSI, may be important?

NATURE GEOSCIENCE DOI: 10.1038/NGEO2640

LETTERS



**Figure 1 | Illustration of photolytic solar cycle ozone response.** The solar max-min photolytic ozone response to integrated-UV below 242 nm (left column) produces ozone. Below 320 nm (middle), O<sub>3</sub> photolysis leads to catalytic loss of O<sub>3</sub> if the O does not recombine with O<sub>2</sub>. The resultant response (right) is approximately the sum of these<sup>9</sup>. The far larger SORCE changes at ozone-destroying wavelengths more than compensates for the larger UV changes at shorter wavelengths, leading to a negative response higher up, whereas SATIRE-S is positive at all altitudes. Pressure levels are illustrative of Fig. 4. Percentages are the relative solar cycle changes.

SIM on SORCE shows, over the period 2003-2007 (i.e. during the declining phase of the most recent solar cycle), a much larger decline in the UV than would be anticipated from current understanding and an *increase* in visible

# Thinkings..

Understanding the role of solar variability in solar activity is essential to the interpretation of past climate and prediction of the future.

Solar activity changes might also play a role in regional climate that we need to understand in the context of informing climate adaptation efforts.

Knowledge of the solar radiation incident on the Earth is an essential prerequisite of any quantitative studies of the Sun's impact on climate (discrepancies of 4-5  $\text{Wm}^{-2}$  between current observations of TSI and estimates of its values back in time to the period of low activity).

Continuing efforts to obtain measurements of high accuracy and precision are needed, alongside further effort in understanding the relationship between activity indicators and irradiance.

# ..Thinkings..

SSI measurements are essential, if SORCE were correct we should revisit our expectations and interpretation of past / future of solar impact on climate

## Mechanisms:

Top-down, uncontroversial. Not however the subsequent impact at the surface

Bottom-up : uncontroversial the first part, not the feedback via the circulation

responses in cloudiness to solar-induced changes in atmospheric ionisation by galactic cosmic rays, requires more evidence

Climate models, need to include all relevant processes..  
However they need correct SSI input data!

# ..Thinkings

## Solar activity and climate change:

Increases in solar activity probably contributed 7-30% of the global warming apparent over the century leading up to the 1960s, the warming in the latter part of the 20th century is almost entirely due to the increasing concentrations of greenhouse gases from human activity

Understanding the physical processes involved in solar-climate connections is crucial to the interpretation of meteorological records, and to the prediction of aspects of the future climate.

Prediction in solar activity are difficult. Moreover, it is not necessarily the case that an Earth with a global net radiation balance but different radiative components (viz. less absorbed solar radiation but more “greenhouse” trapping of infrared radiation) will have the same climate



# Exoplanets exercise

Pos.

Neg.

10.1175/1520-0442(1989)002h0554:ABTYSCi2.0.CO;2.

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