

Beatrice Panico INFN Naples on behalf of PAMELA collaboration

Study on CRE arrival distributions with PAMELA experiment



The PAMELA experiment

Payload for Antimatter/Matter Exploration and Light-nuclei Astrophysics



- Quasi polar and elliptical orbit
- Inclination ~ 70°
- Altitude ~ 300 - 600 km
- From 2010 circular orbit

- Launched on 15th June 2006 - Continuous operation mode

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Searching for anisotropies

Study of e⁺+e⁻ (CREs) arrival directions could help to understand the nature of the positron excess revealed by PAMELA

- > The entire sky can be explored by PAMELA
- Separation of electrons and positrons



Local nearby sources: anisotropy in the angular e⁺e⁻ distribution Dark Matter: the electron/positron pair emission would be strongly correlated with the center of the DM clump and would produce visible effect in anisotropy study.

$\chi\chi$ annihilation γ , e, μ ... γ , e, μ ... γ , e, μ ...

PAMELA lower limit 90% C

10² Energy (GeV)

AMS-02

Adriani et al. Nature 2009

Blind search for anisotropy

Reference map : Arrival directions of particles with an isotropic flux

Event map : Experimental arrival directions The comparison between signal and background to highlight deviations from isotropy is made with different approaches: Li & Ma Statistical significance test Excess in Sun direction? Power Spectrum Analysis

ositr

0.02

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Data sets

Data period : 2006 - 2010~ $1.5 \cdot 10^3$ positrons with energy E > 10 GeV ~ $1.9 \cdot 10^4$ electrons with energy E > 10 GeV ~ $4.1 \cdot 10^5$ protons with energy E > 10 GeV *MC simulation*

~ 10^6 electrons with energy E > 10 GeV

✓ Same period

 \checkmark Selection criteria with same efficiency

✓ Same exposure

Reference frame - Galactic coordinate system The equatorial plane is the projection of the plane of our Galaxy onto the celestial sphere

Visualization of the sky map in a 2D-sphere: **Healpix package**

Each pixel covers the same surface area as every other pixels and is regularly distributed on lines of constant latitudes





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Results

Search in the Sun direction

Some models predict a significant flux of e^+ and e^- around the Sun direction due to DM particles

Method:

- Calculate the events direction in the solar frame
- Calculate the number of events at a given angular distance from the Sun
- Compare the observed events with the expected number of events from an isotropic flux



• Relative map in the basis of spherical harmonics

$$I(l,b) = \sum_{l=1}^{\infty} \sum_{m=-l}^{m=l} a_{lm} Y_{lm}$$

Angular power spectrum obtained

$$C(l) = \frac{1}{2l+1} \sum_{m=-l}^{m=l} a_{lm}^2$$

• Dipole upper limit δ = 0.166 at 95% C.L.

