

International School of Space Science (ISSI)

*The Polar Upper Atmosphere: From Science to
Operational Issues, L'Aquila, Italy, 16-21 Sep. 2018*

**“MAKING THE IONOSPHERE
— All in a Morning’s Work”**

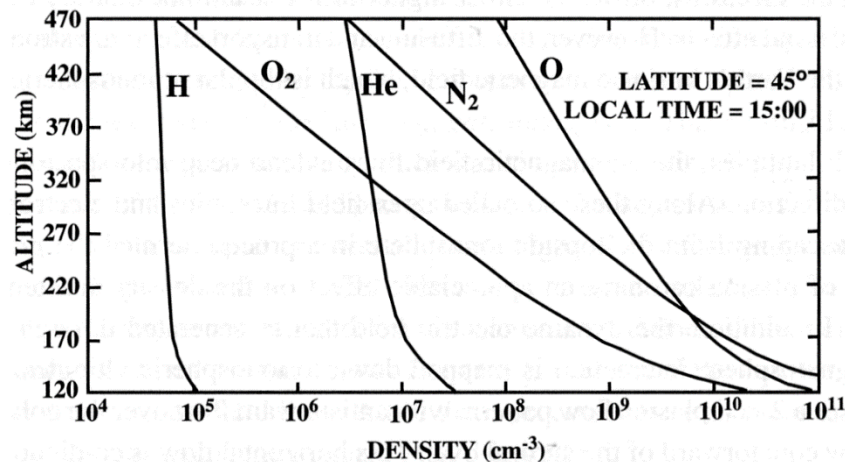


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1. The Photo-Ionization Process

- Start with Neutral Atmosphere



- Ionization Potentials of Atoms and Molecules

Species	Energy (ev)
O	13.62
O ₂	12.06
N ₂	15.58

$\lambda_{\text{ionization}}$
910 Å
1028 Å
796 Å

- Photon energy, $E = h\nu = hc/\lambda$

$$\text{or } \lambda(\text{\AA}) \approx \frac{12345}{E(\text{ev})}$$

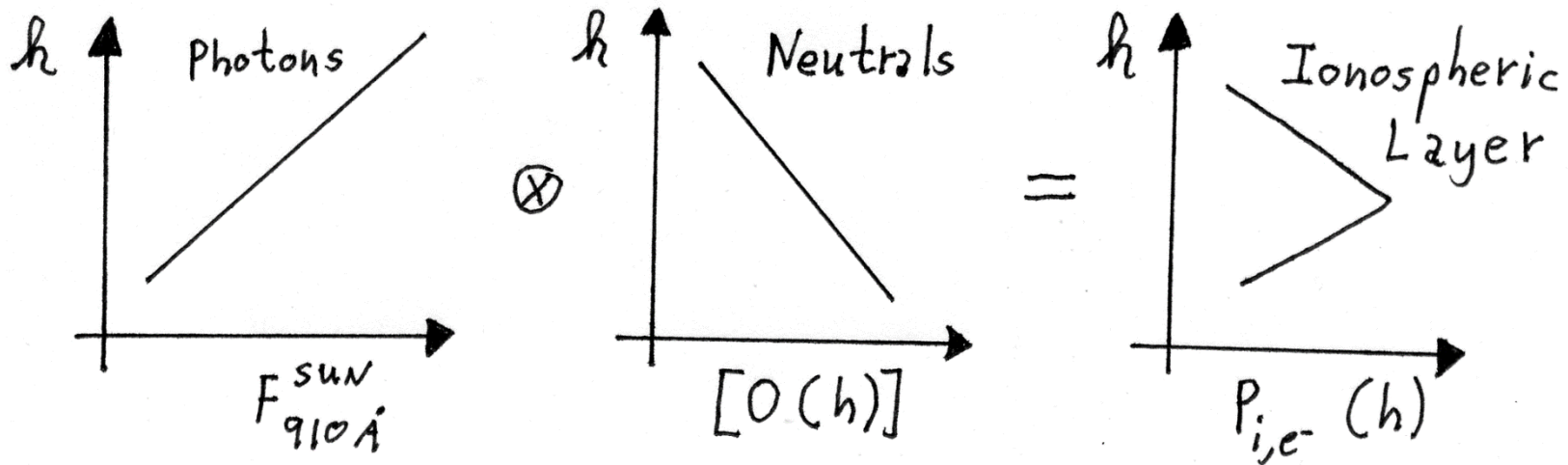


Thus, Photon (910 \AA) + O \rightarrow O $^+$ + e $^-$

Knowing $F_{910}^{\text{SUN}}(h) \text{ \& } [O(h)] \rightarrow [O^+(h)] + [e^-(h)]$

Production Function (P)
for monochromatic ionizing radiation
(called "Chapman Theory")

How should P(h) look?

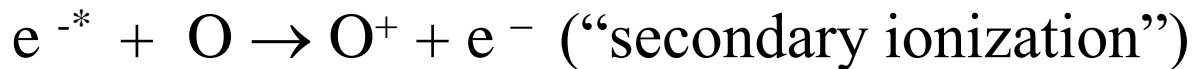


Algebraically $P_{i,e^-}(h) = F_{\lambda_{ion}}^{sun}(h) \cdot \underbrace{\sigma_{ion}}_{\substack{\text{ionization} \\ \text{cross-section}}} \cdot [O(h)]$

What happens if photon $\begin{cases} \lambda > \lambda_{ion} ? \rightarrow \text{No ionization} \\ \lambda < \lambda_{ion} ? \rightarrow \text{Extra Energy} \end{cases}$

i.e., Photon ($\lambda < \lambda_{ion}$) + O \rightarrow O⁺ + e^{-*} (energetic photo-electron)

If e^{-*} Kinetic Energy > Ionization Potential
collisions can cause additional ionizations



Thus, a very energetic photon can lead to several ion-electron pairs.

For a complete model of Photo-Ionization, the flux of solar photons at all relevant λ s is needed:

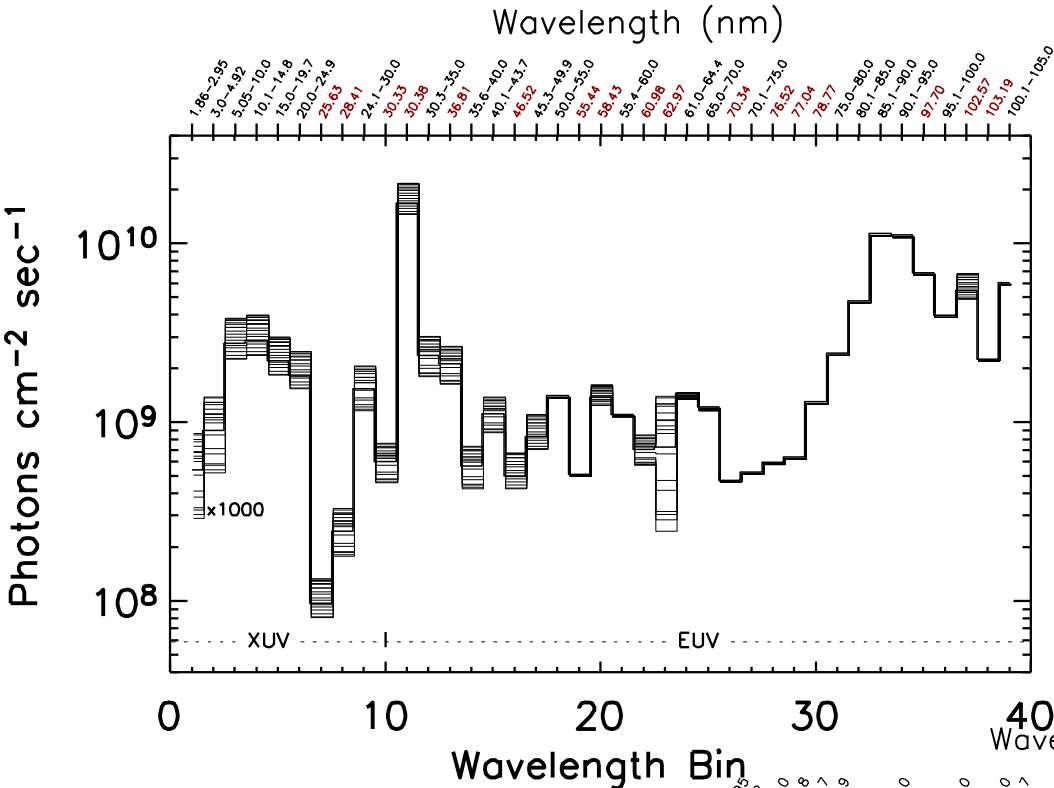
$$P_{\text{Total}}(h) = \sum_{\lambda=0}^{\lambda_{\text{ion}}} F_{\lambda}^{\text{sun}}(h) \cdot \sigma_{\text{ion}}(\lambda) \cdot [N(h)]$$

Table 2.1. *Solar spectral regions.*

Radio	$\lambda > 1 \text{ mm}$
Far Infrared	$10 \mu\text{m} < \lambda < 1 \text{ mm}$
Infrared	$0.75 \mu\text{m} < \lambda < 10 \mu\text{m}$
Visible	$0.3 \mu\text{m} < \lambda < 0.75 \mu\text{m}$
Ultraviolet (UV)	$1200 \text{ \AA} < \lambda < 3000 \text{ \AA}$
Extreme ultraviolet (EUV)	$100 \text{ \AA} < \lambda < 1200 \text{ \AA}$
Soft x-rays	$1 \text{ \AA} < \lambda < 100 \text{ \AA}$
Hard x-rays	$\lambda < 1 \text{ \AA}$

Note: $\text{\AA} = 10^{-10} \text{ m}$.

(1 $\text{\AA} = 0.1 \text{ nm}$)



SOLAR2000 IRRADIANCES

* Flux in 39 wavelength bins (bands + **lines**)

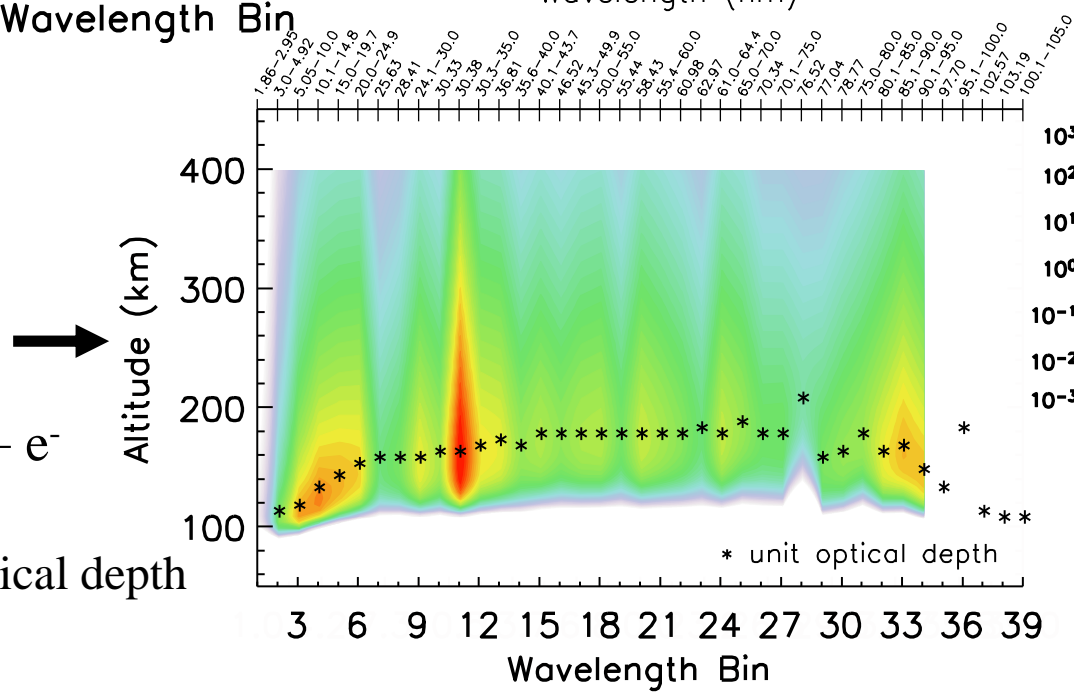
* Shown here are for 15 consecutive days to portray variability

Sample results:

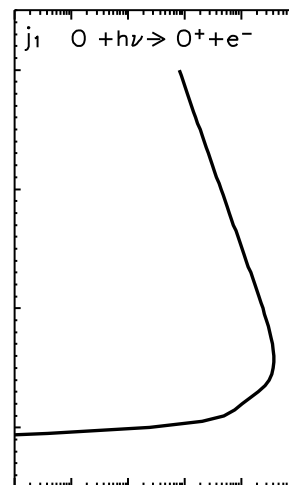
Oxygen

Photons + $\text{O} \rightarrow \text{O}^+ + \text{e}^-$

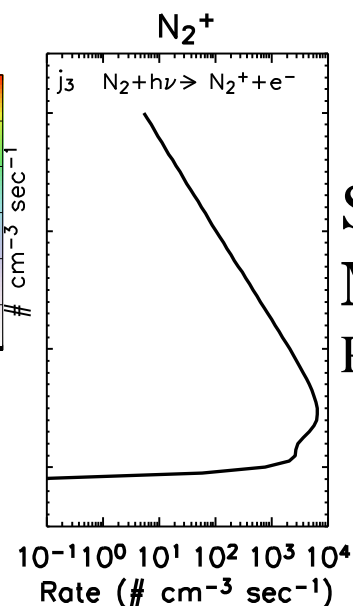
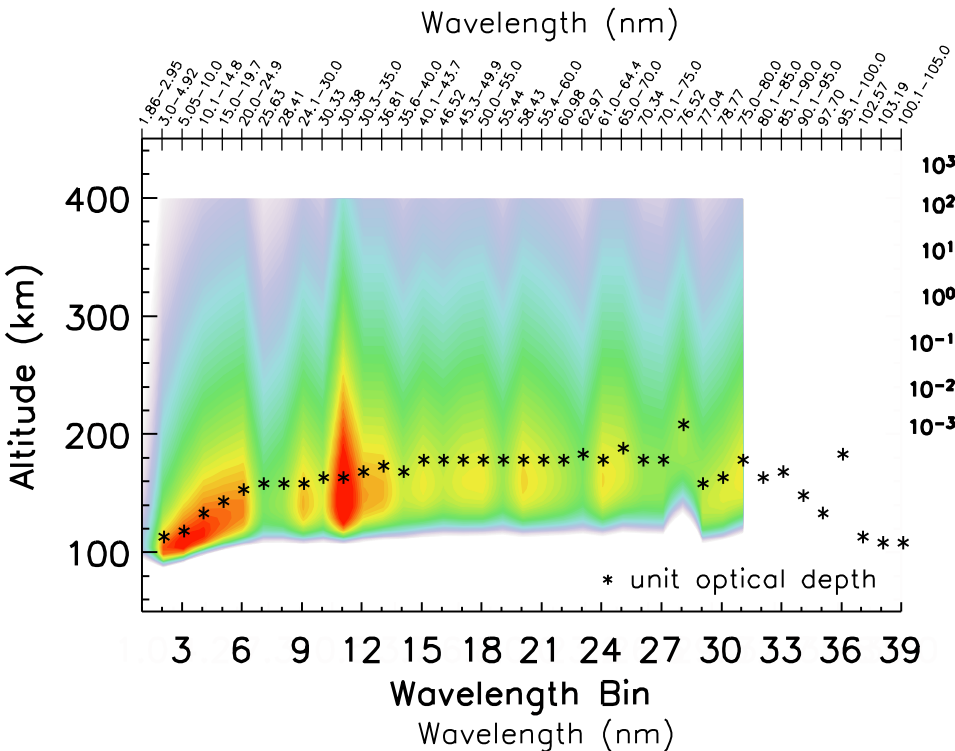
* = unit optical depth



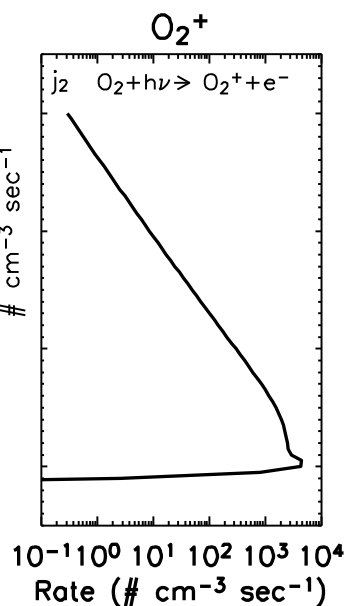
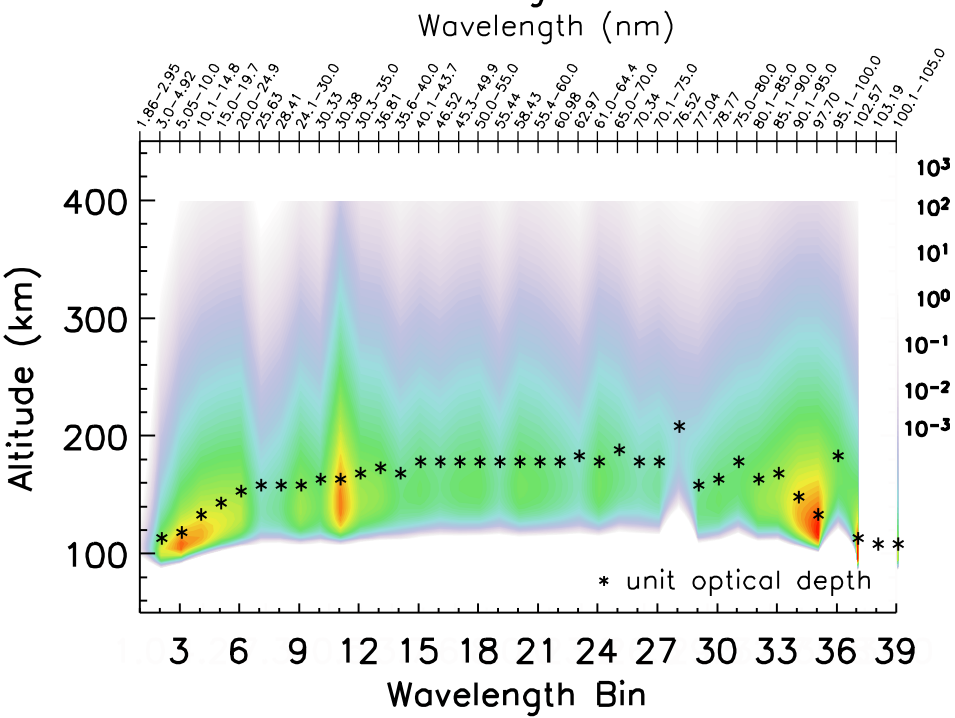
O^+



Rate (# $\text{cm}^{-3} \text{sec}^{-1}$)



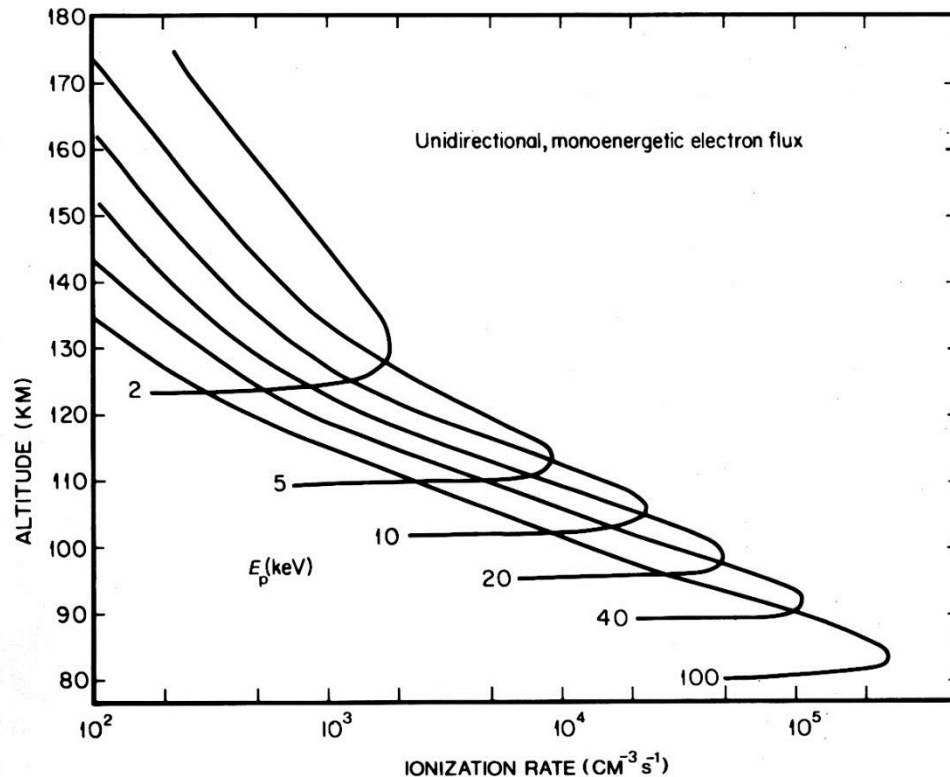
Sample results:
Molecular Nitrogen
 Photons + $N_2 \rightarrow N_2^+ + e^-$



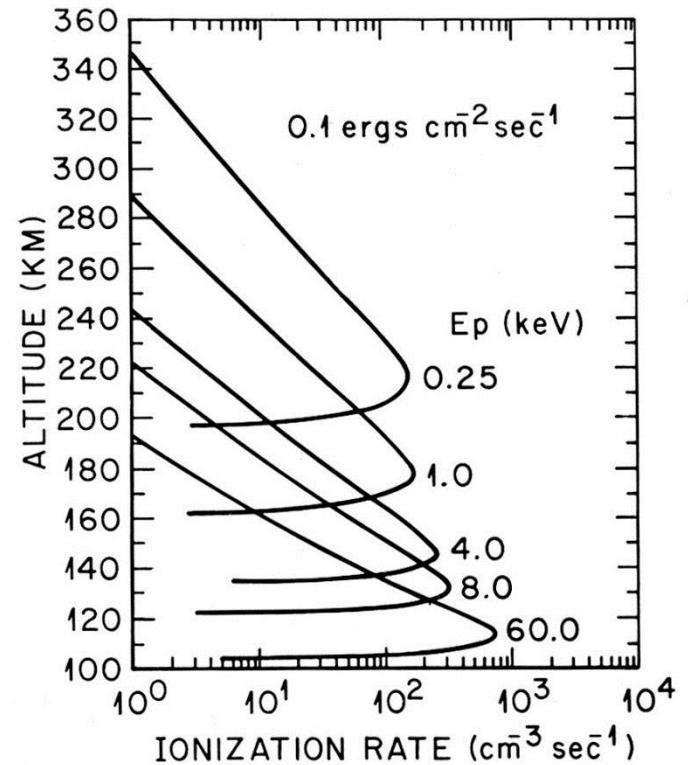
Sample results:
Molecular Oxygen
 Photons + $N_2 \rightarrow O_2^+ + e^-$

Production of Ionospheric Plasma by Energetic Particles

Precipitating Electrons



Precipitating Protons



--- taken from Rees (1989)

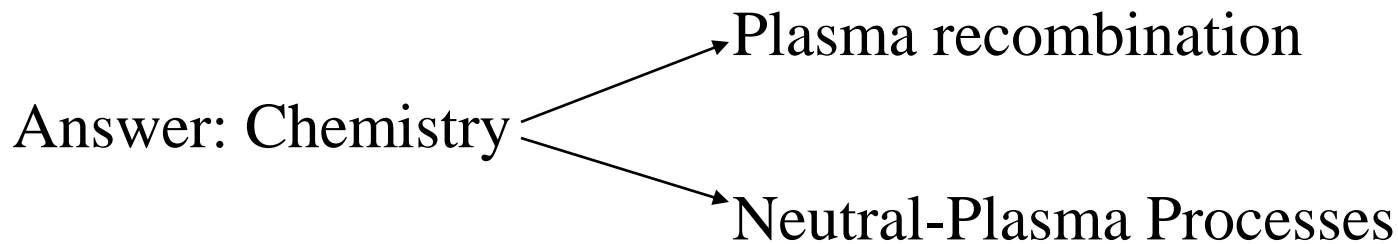
2. Ionospheric Transformations

- What does “production only” imply?
e.g., use $P(O^+)$ value from graphs (photons or particles)

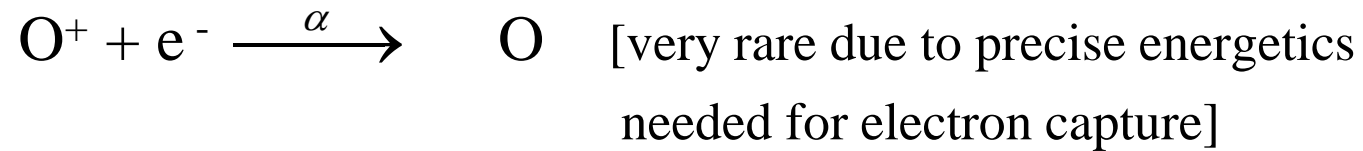
$$P_{\max} = 4000 \text{ e}^-/\text{cm}^3/\text{sec} \times 3 \text{ hours } (\approx 10^4 \text{ sec})$$

$$\text{gives } N_{\max} \approx 4 \times 10^7 \text{ e}^-/\text{cm}^3 \quad \textit{Never Measured!!!}$$

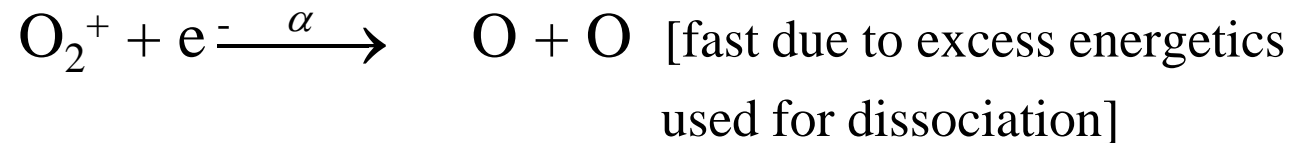
Message: Something happens to these ions and electrons



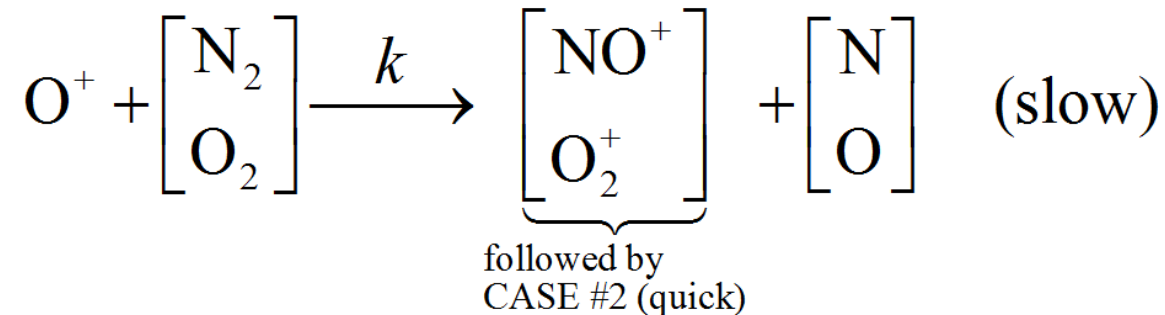
CASE # 1: Atomic ions + electrons



CASE # 2: Molecular ions + electrons



CASE #3: Transform Atomic ions to Molecular ions

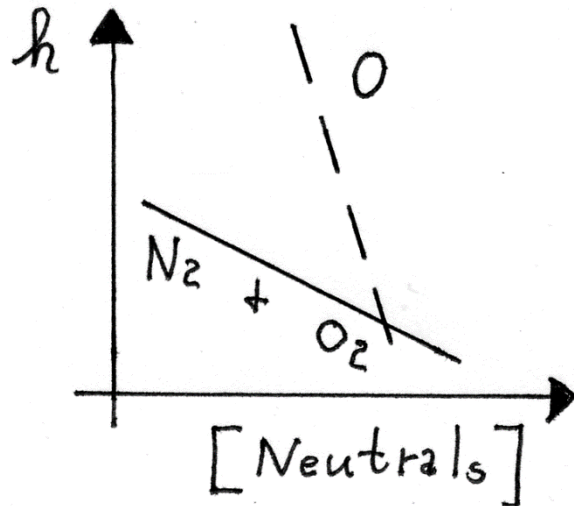


The 2-stage recombination process governed by slower step, e.g.,

$$\boxed{\frac{dN_e}{dt} = -k[\text{N}_2]N_e = -\beta N_e}$$

Messages from Simple Photochemical Theory

- Plasmas should be ionized form of dominant neutral



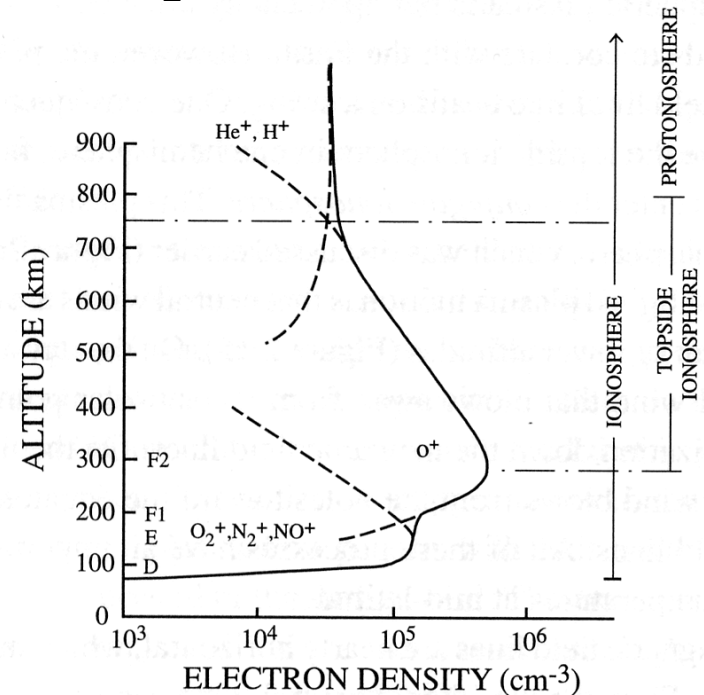
→ CASE #1 : $O^+ + e^-$

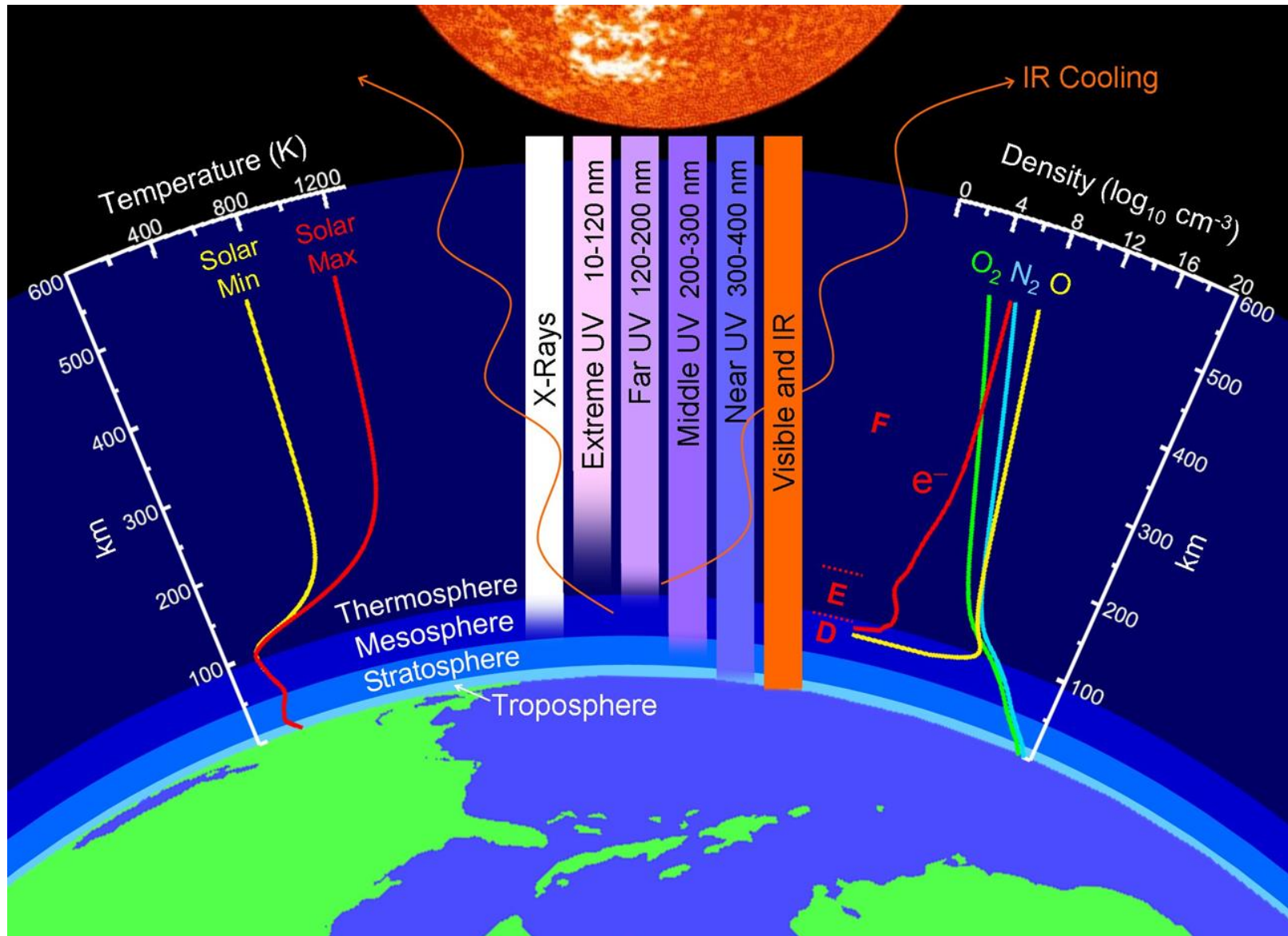
→ CASE #2 : $O_2^+ + e^-$
 $N_2^+ + e^-$

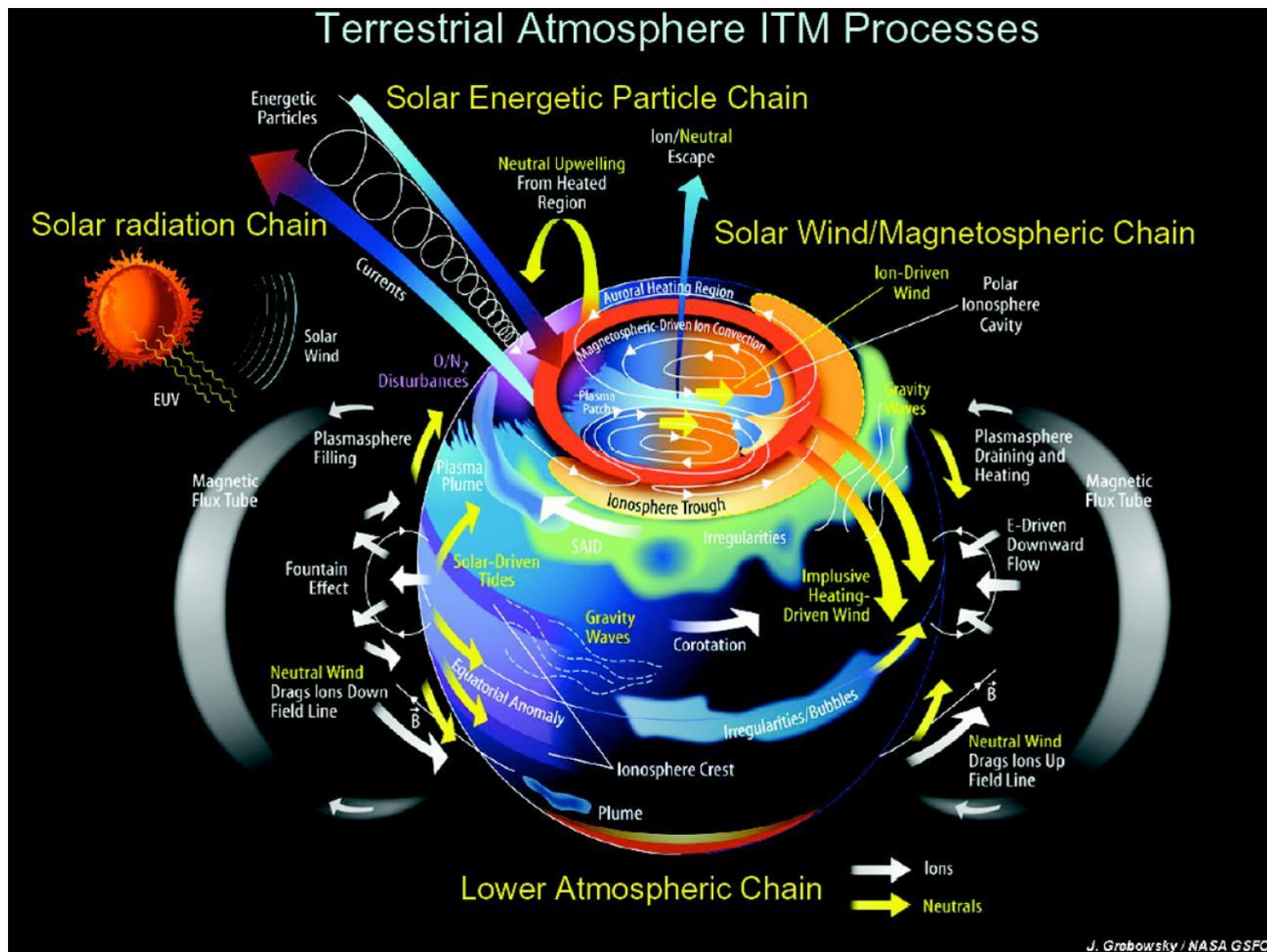
- The actual case:

— some chemical transformations
 to form NO^+ and H^+

- Two main layers: F-layer (EUV) and E-layer (X-rays)

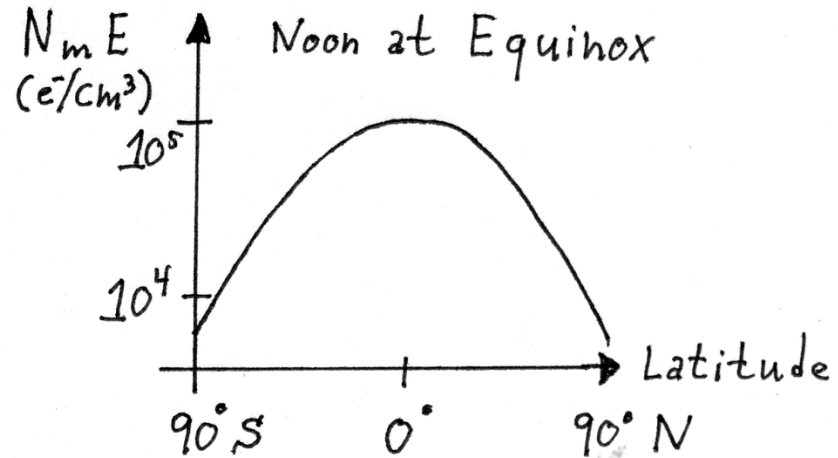
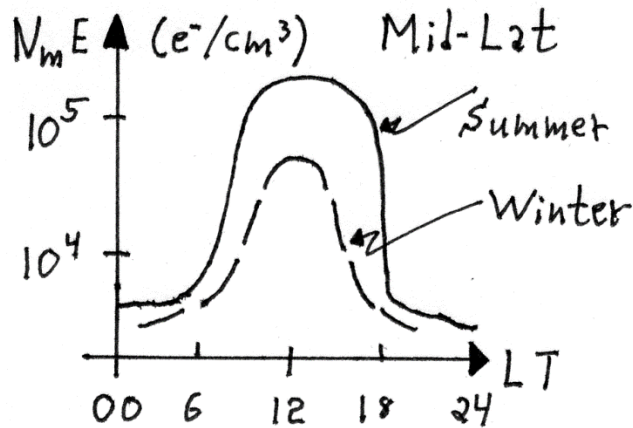






Some E-layer Characteristics

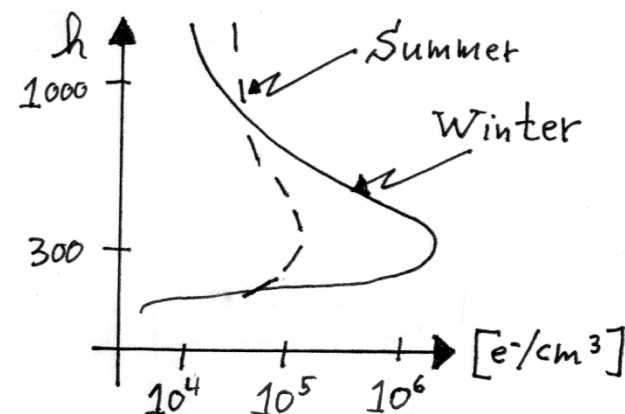
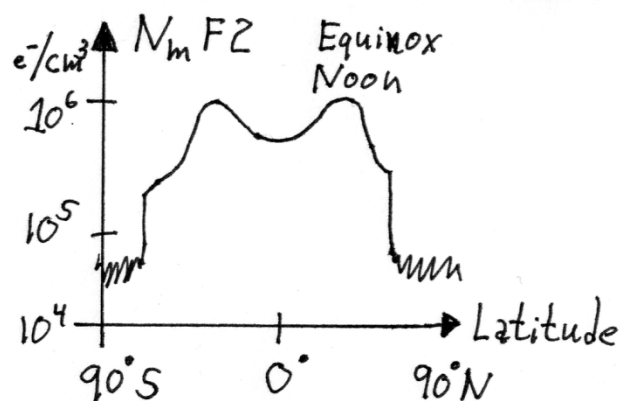
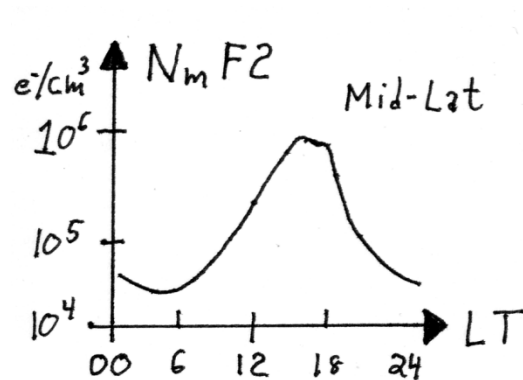
- In regions of a dense neutral atmosphere ($h \leq 150$ km) all ions are molecular (rapid chemistry) and the ions + electrons stay where produced (too many collisions to move away).
- Example of diurnal behavior



The E-layer is controlled by the Sun's flux and its position ($\text{dec} + \chi_\odot$)

3. Photochemistry-Plus-Dynamics

- Some F-layer Characteristics



The F-layer is produced by sunlight BUT its behavior does not follow $\chi_\odot \Rightarrow$ “Anomalies”

What are the causes of vertical motions?

- GRAVITY

- For neutral gas, $\frac{dP}{dh} = -\rho g \Rightarrow \frac{d(NkT)}{dh} = -Nmg$

Solution: Hydrostatic Law

$$N(h) = N_o(h_o)\exp[-(h-h_o)/H_n]$$

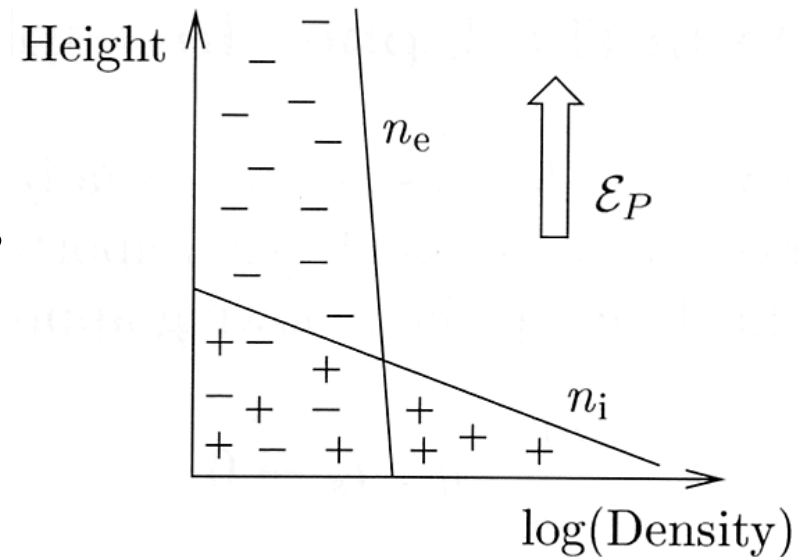
neutral scale height $H_n = \frac{kT}{mg}$

- For plasma \rightarrow ions + electrons

$$m_i \gg m_e$$

Gravity tends towards “charge separation”

\Rightarrow Polarization \vec{E} -field (ϵ_p)



$$\frac{dP_i}{dh} = -\rho_i g + N_i e \epsilon_p$$

$$\frac{dP_e}{dh} = -\rho_e g - N_e e \epsilon_p$$

Adding, with $N_i = N_e = N$ and $m_e \leftrightarrow 0$

$$\frac{d(P_i + P_e)}{dh} \approx -Nm_i g$$

- Solution $N(h) = N_o(h_o)\exp[-(h-h_o)/H_p]$

Plasma scale height $H_p = \frac{k(T_e + T_i)}{mg}$

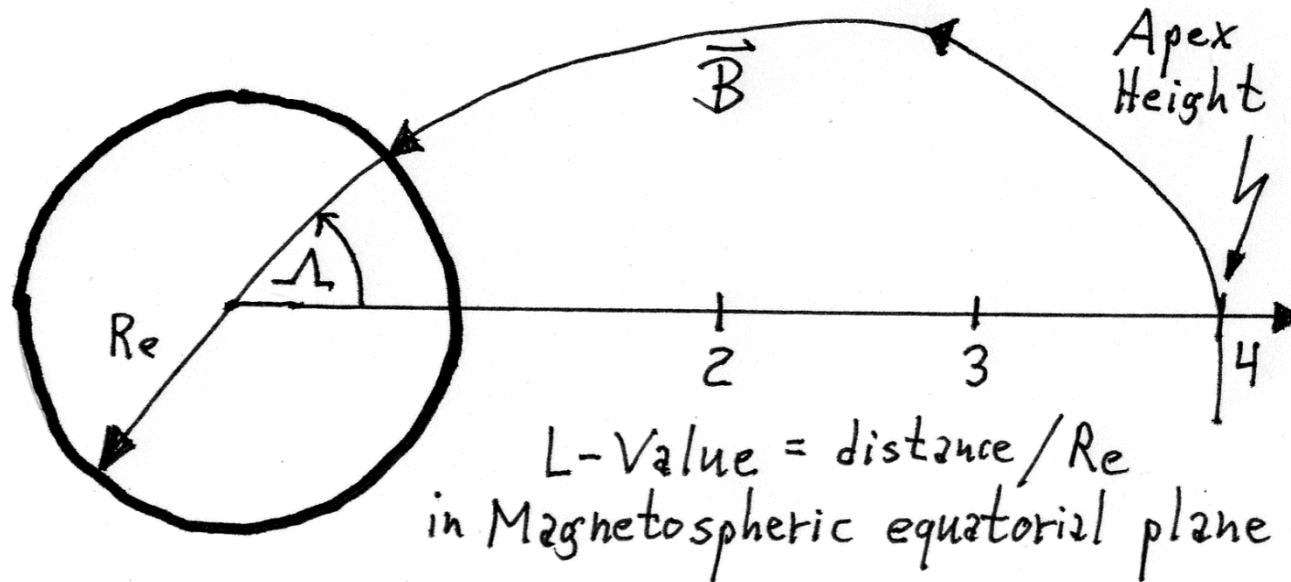
$$H_p \geq 2H_n$$

Called “Ambipolar diffusion”

electrons “pull” ions upward

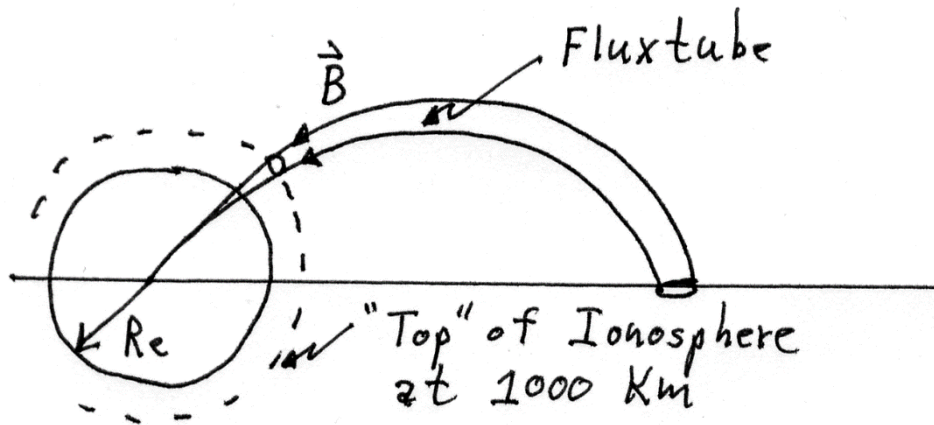
ions “pull” electrons downward

- Midlatitudes (extended)



Magnetic Latitude (Λ), $\cos^2 \Lambda = 1/L$ (e.g., $L=4 \Rightarrow \Lambda = 60^\circ$)

What happens to ionospheric plasma diffusing upwards?

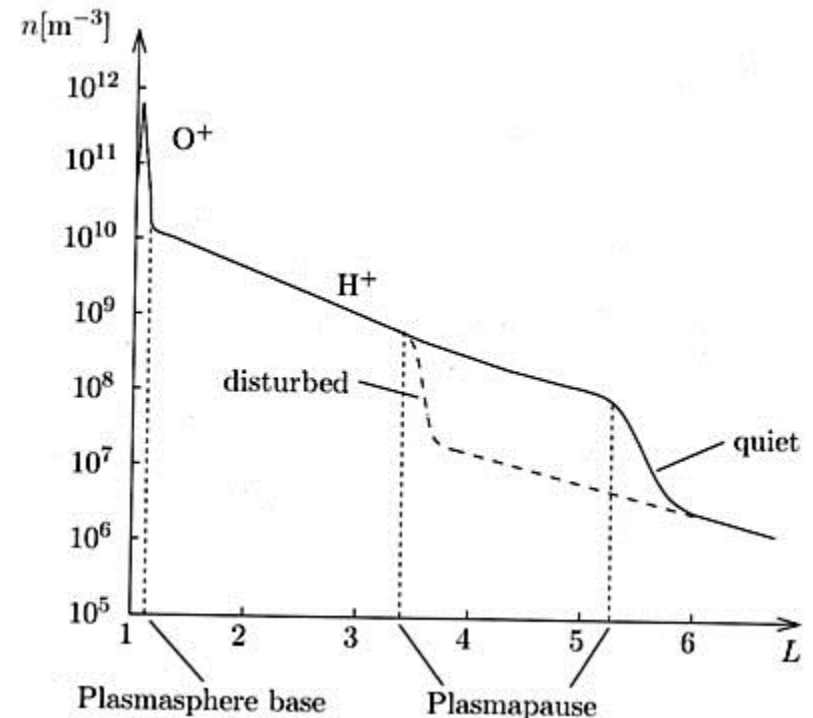
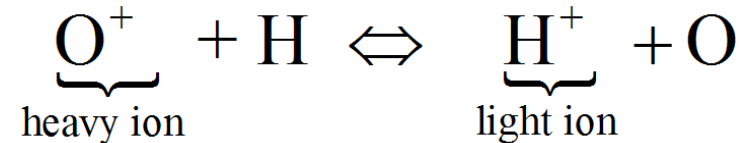


Diffusion affected by reduced gravity, $F_{\text{centripetal}}$, light ions

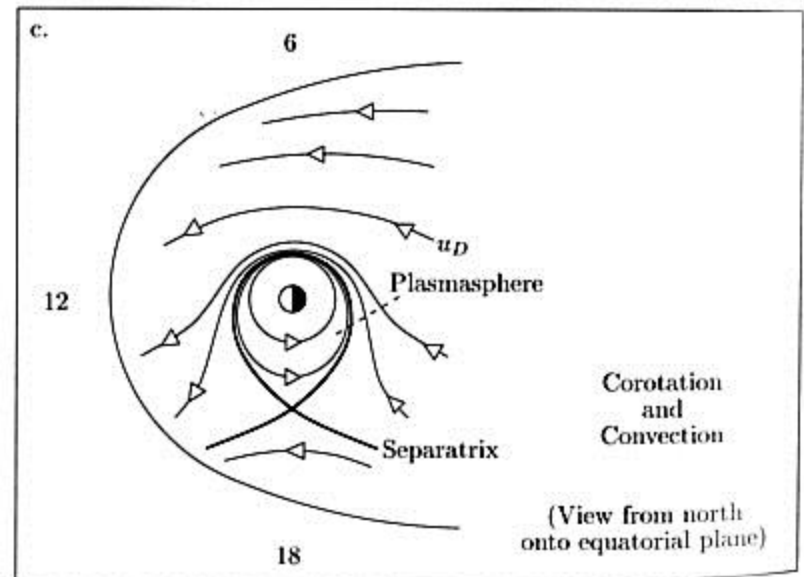
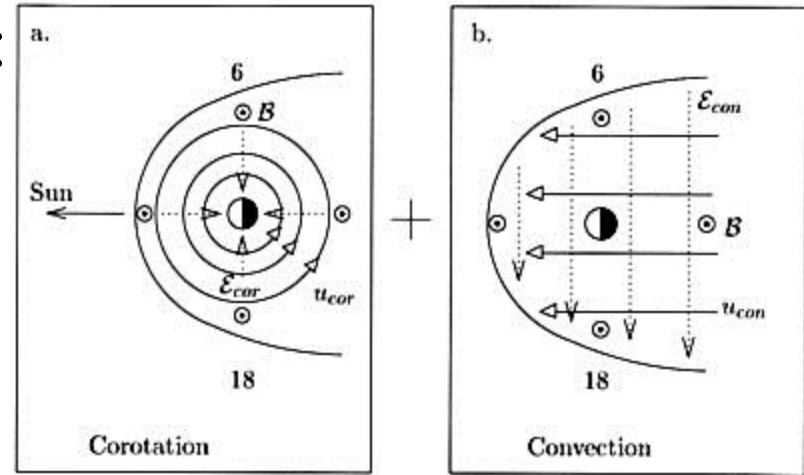
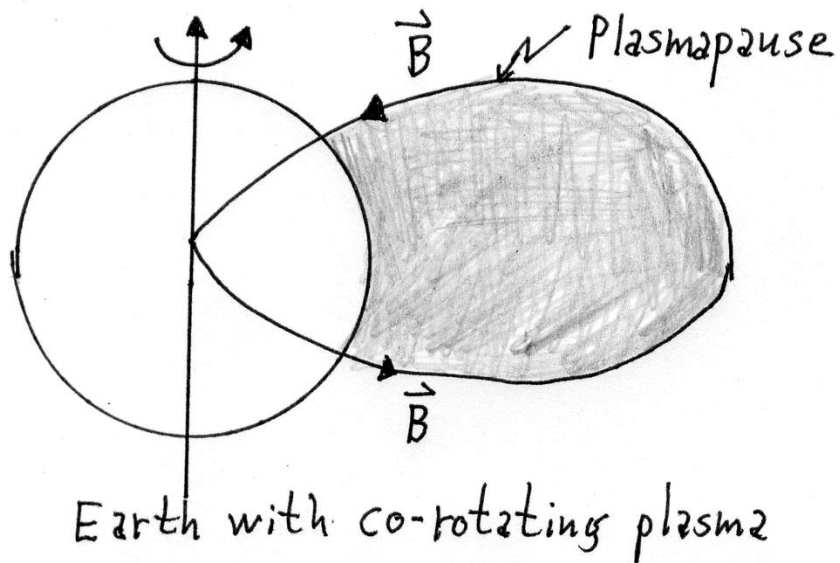
$$\frac{dP}{ds} \rightarrow \text{large } H_p$$

→ fluxtubes of plasma of ionospheric origin

"Charge Exchange"

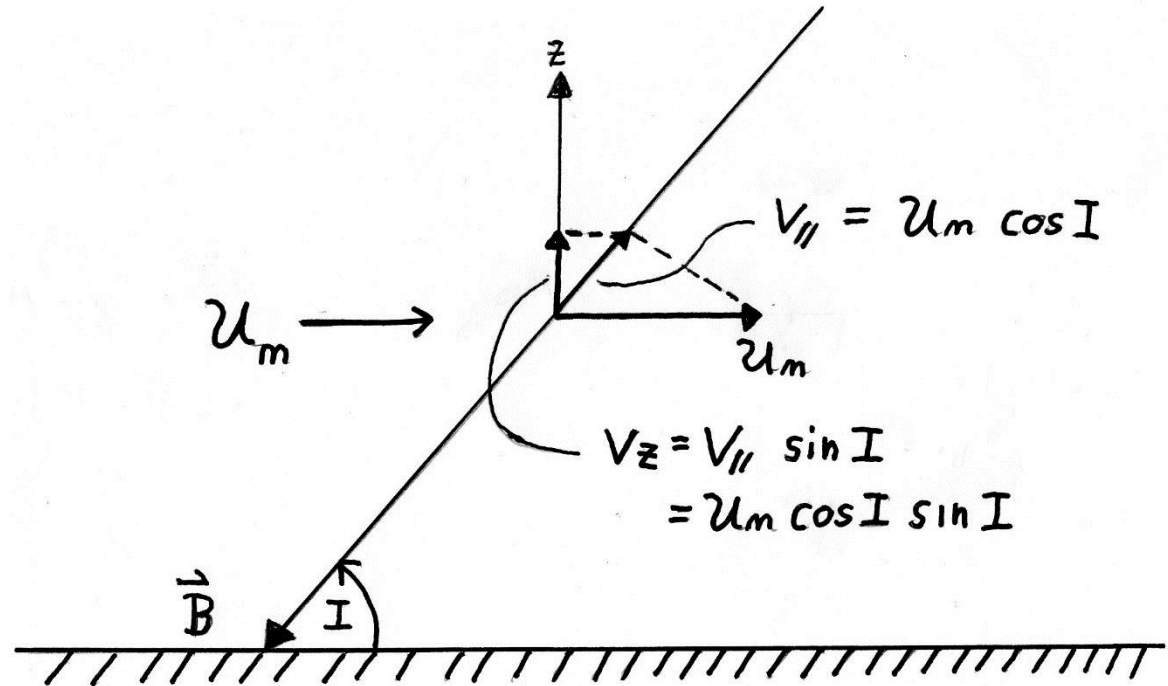


The Plasmasphere in 3-dimensions:



What else causes Vertical Motions? Roles of Magnetic Field

- Neutral Winds (U_m) are horizontal
- Plasma constrained to move $\parallel \vec{B}$

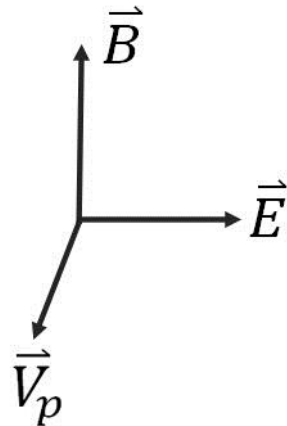


- Middle Latitudes – maximum effect
 - Equatorial Latitudes ($I = 0^\circ$) – small effect
 - High Latitudes ($I = 90^\circ$) – small effect
- Unless U_m generates polarization \vec{E} -field

Electrodynamics: Motions caused by induced or penetrating \vec{E} -fields

$$\vec{V}_p = \frac{\vec{E} \times \vec{B}}{|\vec{B}|^2}$$

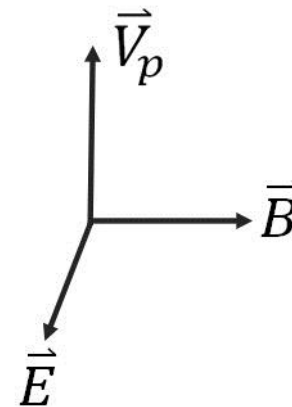
At high latitudes



\vec{B} close to vertical
 \vec{E} horizontal causes
 horizontal \vec{V}_p

Convection Patterns

At low latitudes



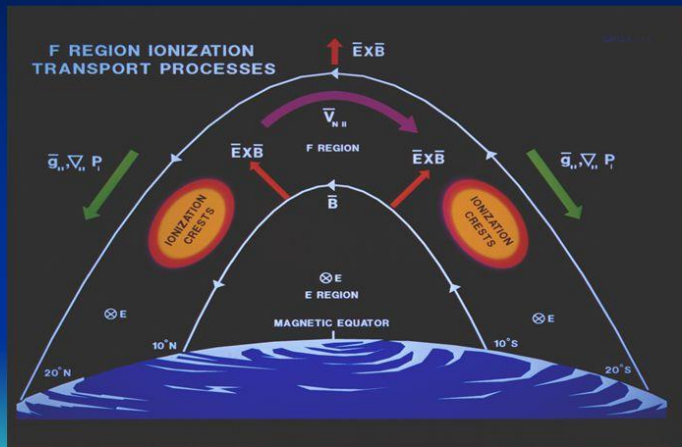
\vec{B} close to horizontal
 \vec{E} horizontal causes
 vertical \vec{V}_p

Fountain Effect

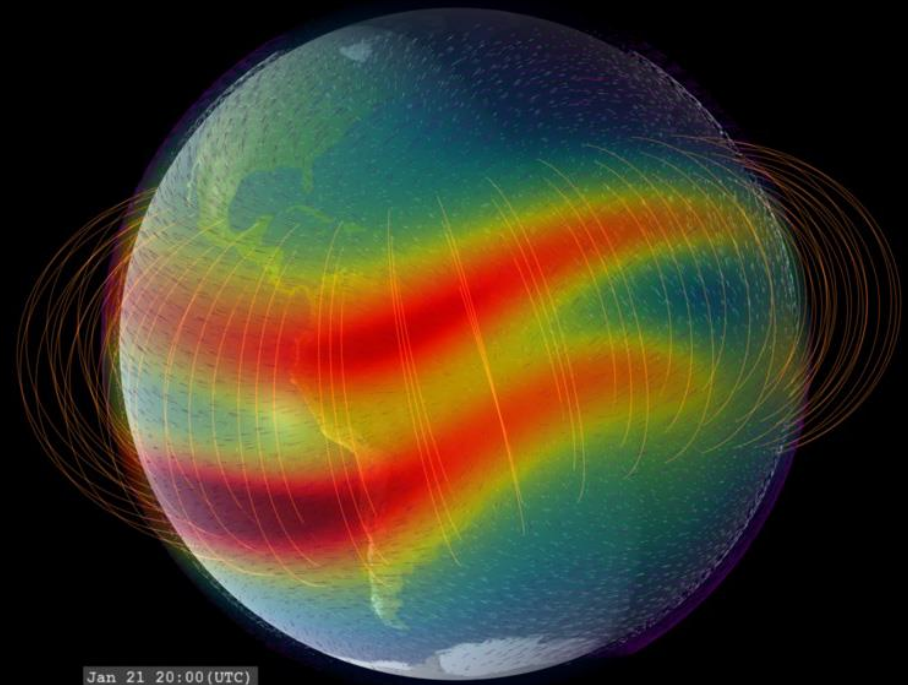
Effects of Electro-Dynamics at Low Latitudes

Equatorial Plasma Fountain & The Equatorial Ionization Anomaly (EIA)

Low-Latitude Ionosphere is dominated by the electrodynamics



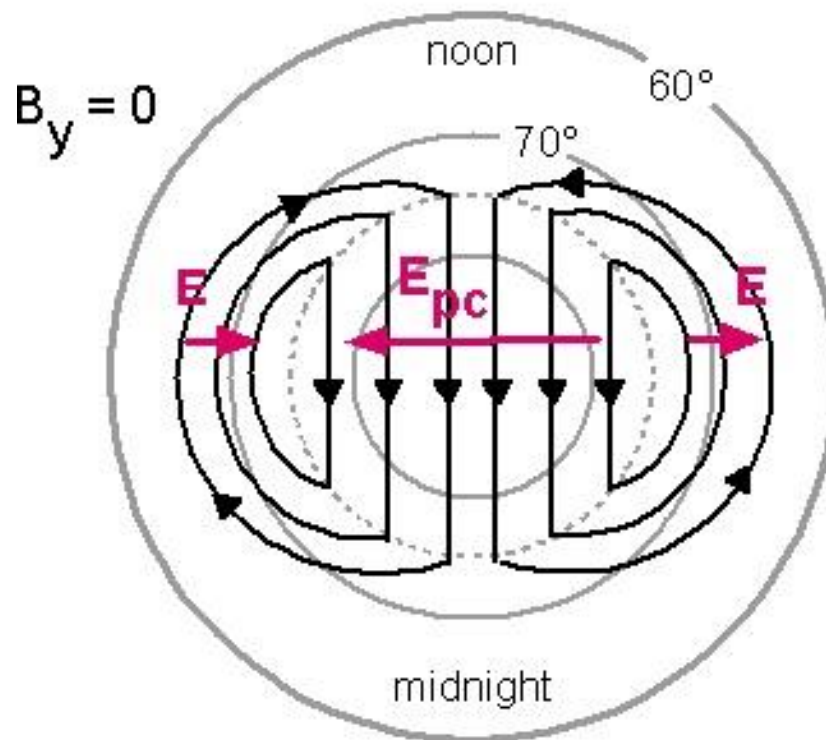
Courtesy of David Anderson



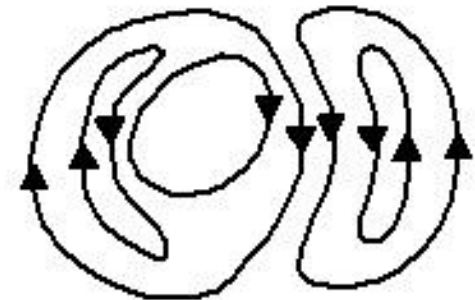
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Effects of Electro-Dynamics at High Latitudes

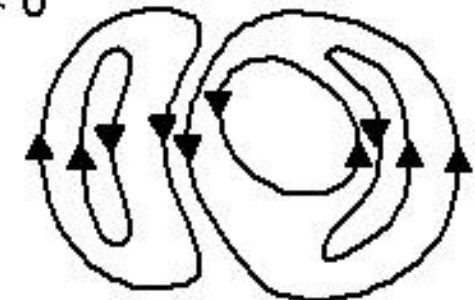
B_z southward



$B_y > 0$



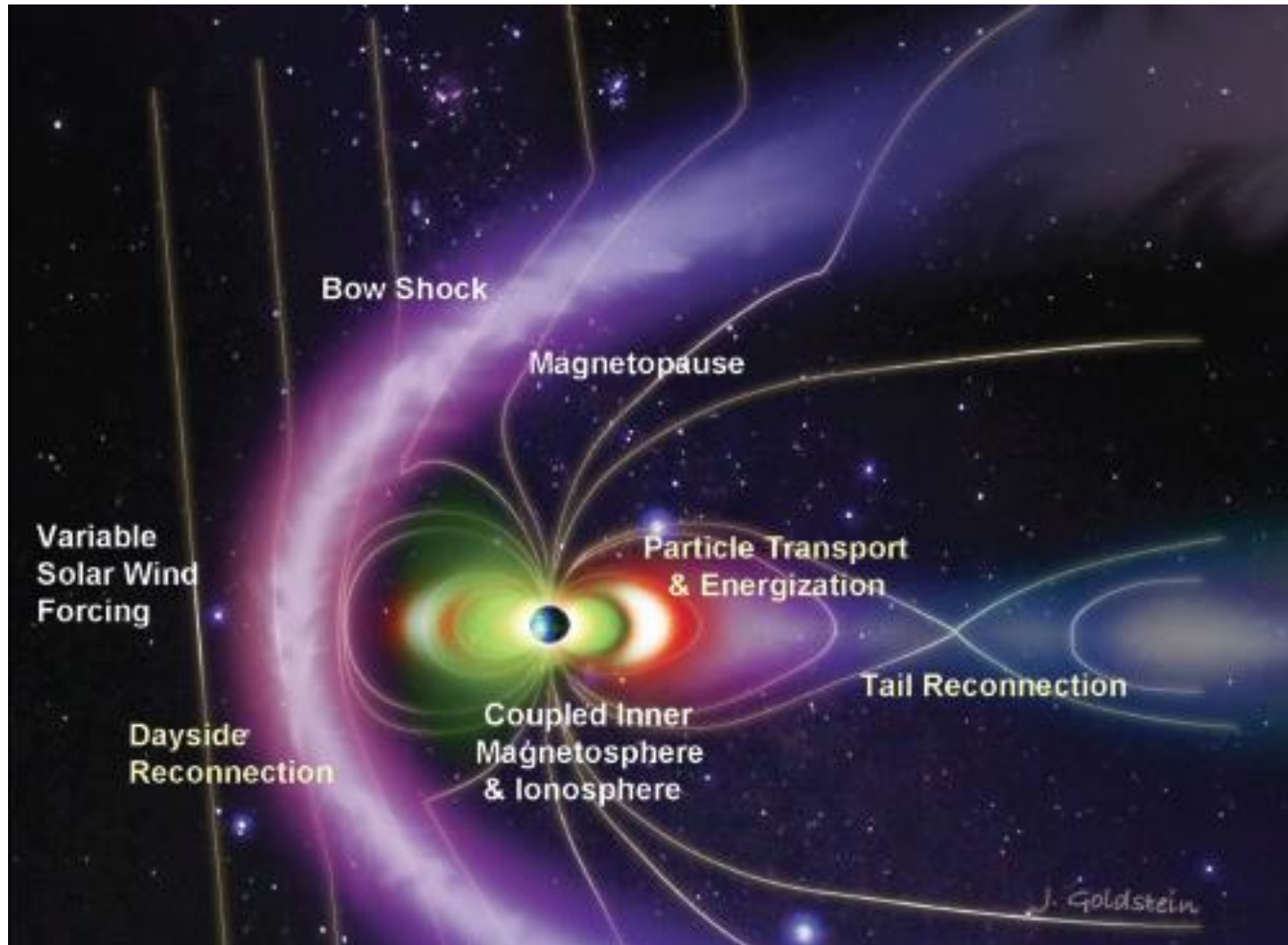
$B_y < 0$



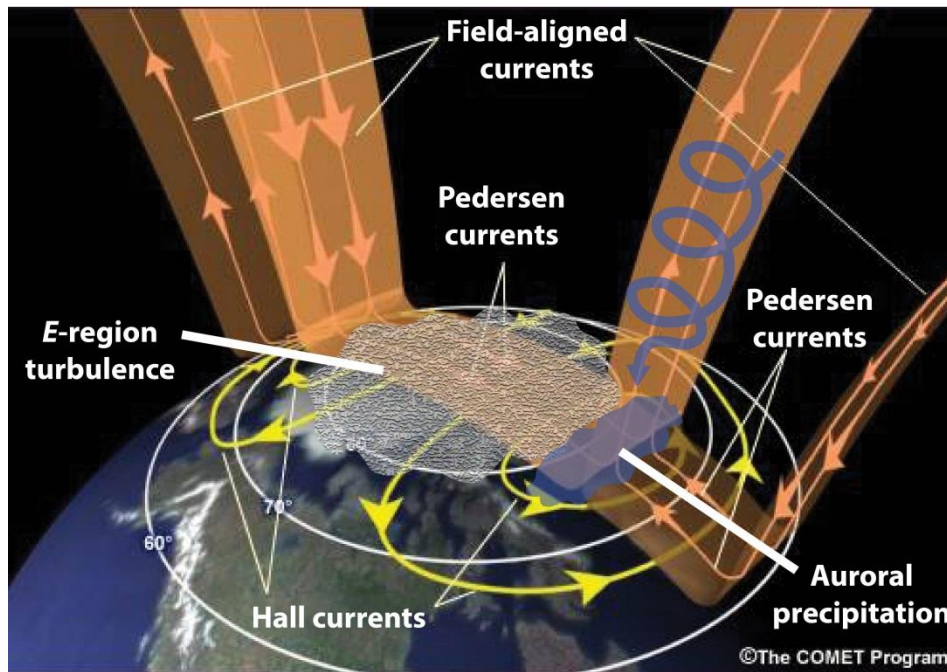
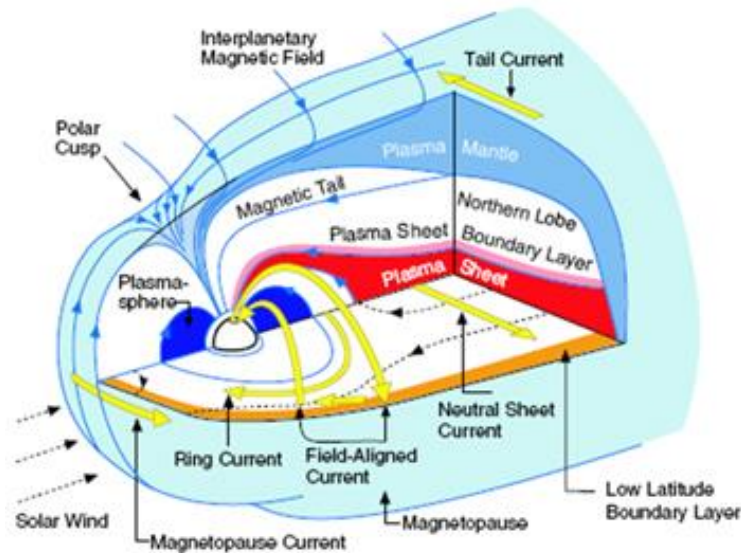
The Polar and High Latitude Ionosphere

- Minimal production of plasma by solar photons
- Strong production of plasma by precipitation of energetic particles from Magnetosphere (nightside) and solar wind (dayside)
- Strong \vec{E} -fields:
Solar wind \rightarrow Magnetosphere \rightarrow Ionosphere
- Electrodynamics: Horizontal convection/circulation patterns
- Upward diffusion & escape: Polar Wind
- Many ionospheric irregularities due to Plasma Instabilities

Summary #1: External Drivers



Summary #2: A Highly Coupled/Complex System



Current closure (height-integrated Ohm's law) is **key to M-I-T coupling**

$$\nabla_{\perp} \cdot (\vec{\Sigma} \cdot \nabla_{\perp} \Phi) = J_{\parallel}$$

$$\vec{\Sigma} = \int \vec{\sigma} dh$$

Conductivity regulates interaction between the field-aligned current J_{\parallel} , electric potential Φ and Hall and Pedersen currents.

Conductivity $\vec{\sigma}(n, n_e, n_i)$ and its functional dependences are **critical**.

Need global distributions of neutral (n), electron (n_e), ion (n_i) densities.

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